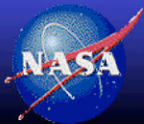

CERES Edition 3 Test Calibration Theory and SSF Results

**CERES Instrument Working Group
CERES Science Team Meeting,
Newport News, VA, 04/24/07**

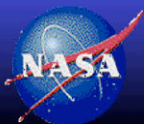
Contributors: Norman Loeb, Costy Loukachine, Bruce Wielicki



NASA Langley Research Center / Science Directorate

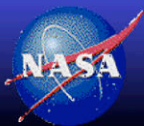


-
- Recap on a new climate ERB calibration stability goal and Rev1 Adjustment Results for Terra and Aqua
 - Detecting illusive spectral darkening: Stability metrics of DCC albedo and nadir direct compare
 - Improved SW darkening model and application to Total channels
 - Results of Terra & Aqua Edition 3 Test calibration through SSF inversion and compared to Edition 2
 - Summary and thoughts for commentary

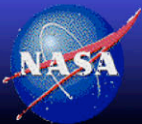
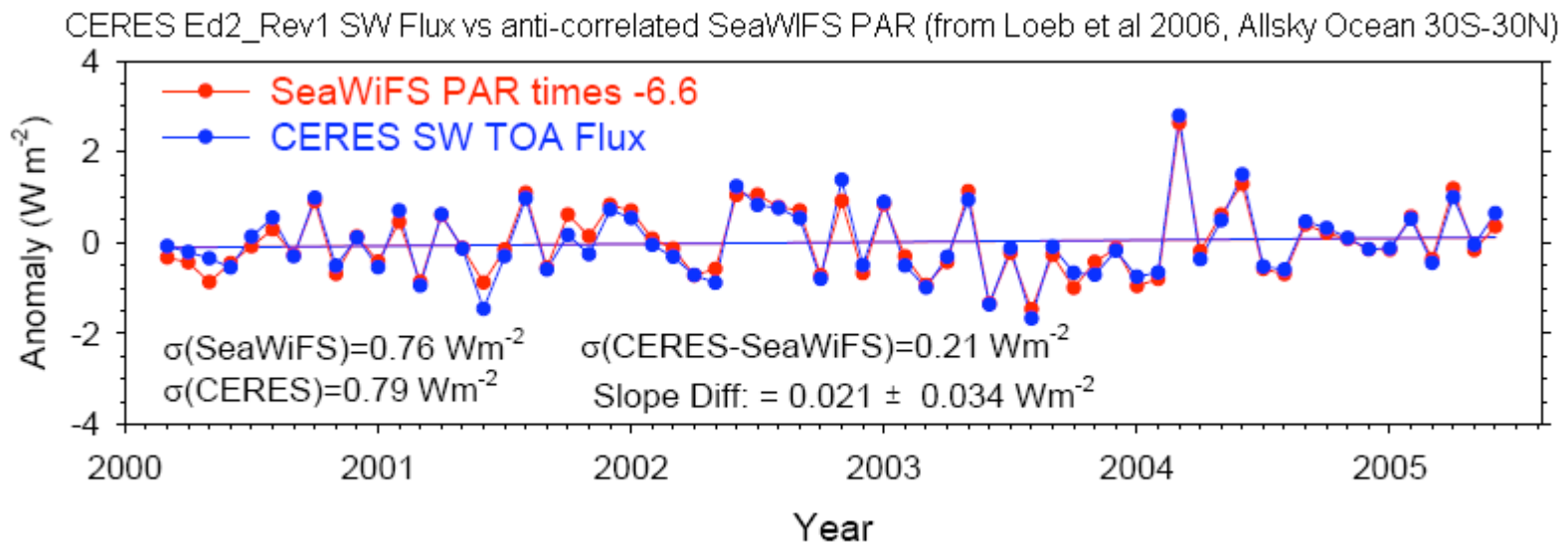


Ohring et al (2005)
suggests that ERB
measurement stability
needs to reach 0.3% per
decade

This is near to an order of
magnitude greater than
CERES instrument design
specification stability
(note: not data products)



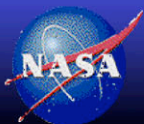
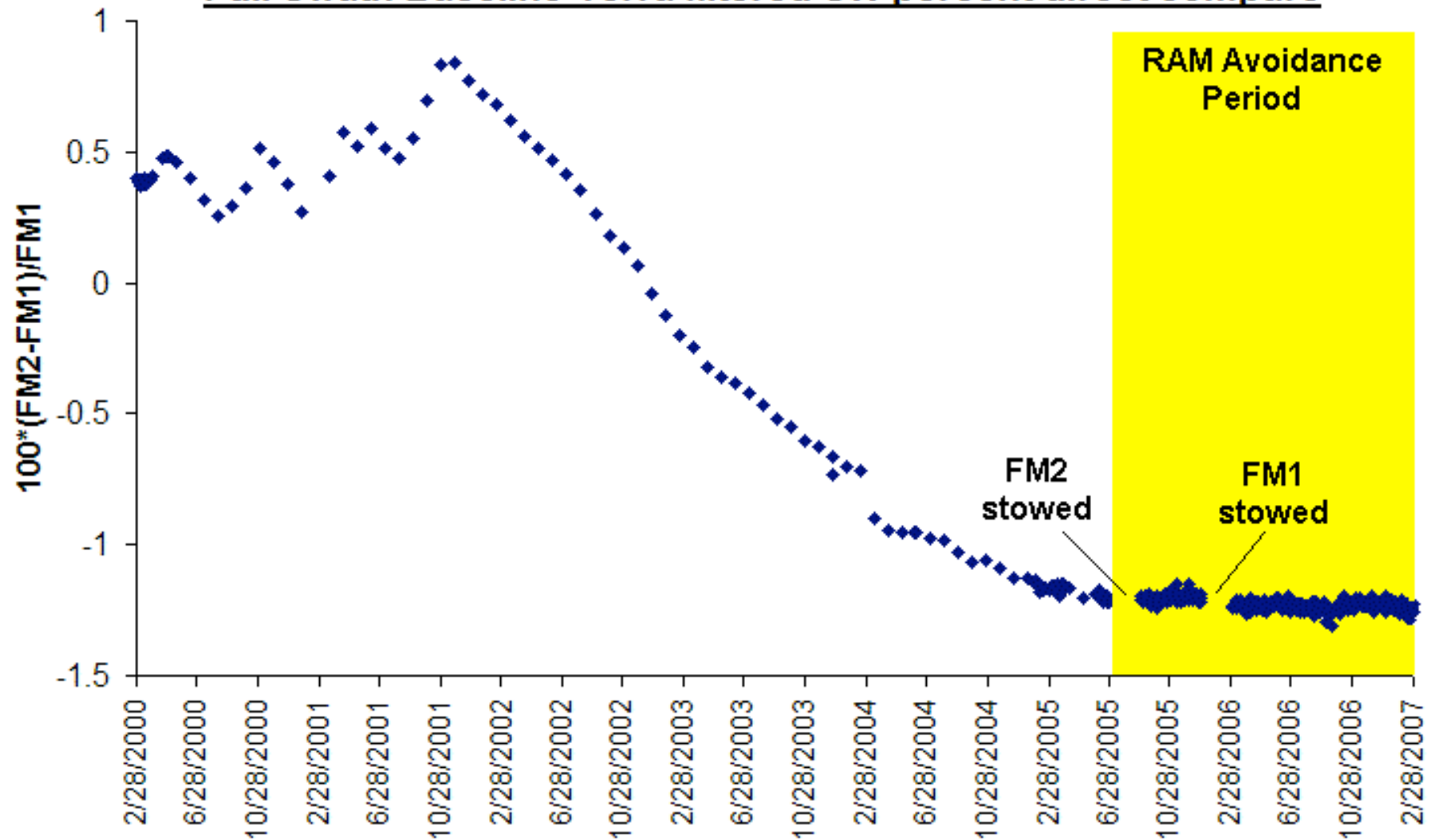
The resulting comparison with anti-correlated SeaWiFS PAR shows that the Rev1 adjustment to CERES Terra SW data yields stability comparable to that of SeaWiFS:



NASA Langley Research Center / Science Directorate



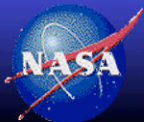
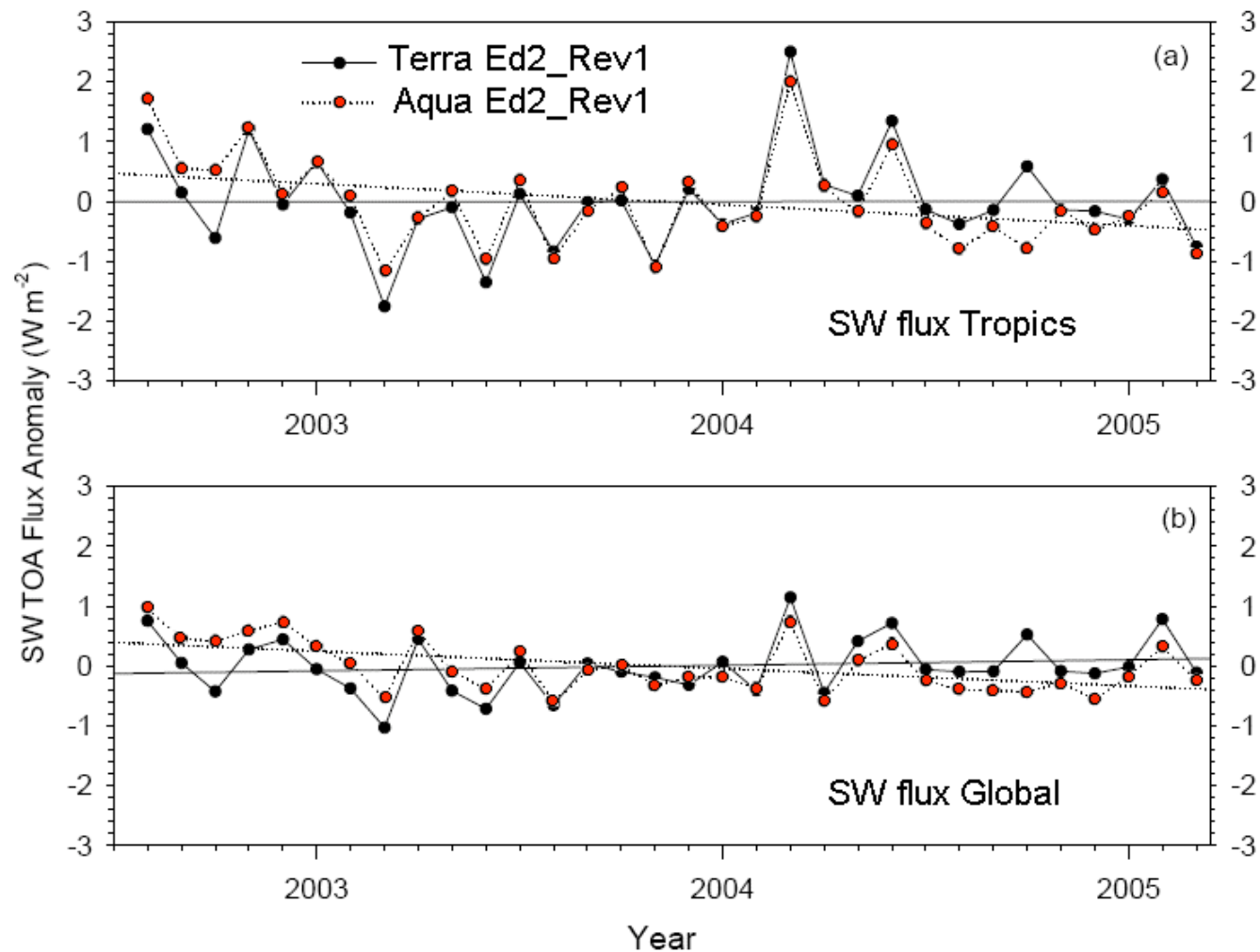
Full Swath Baseline Terra filtered SW percent direct compare



NASA Langley Research Center / Science Directorate



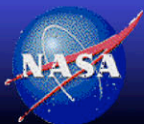
However, even after the Rev1 adjustment the Aqua SW flux anomaly continues to drop relative to Terra:



Examples of Edition 2 Metrics for Stability

1. Edition 2 Deep Convective Cloud Albedo

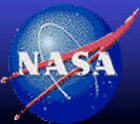
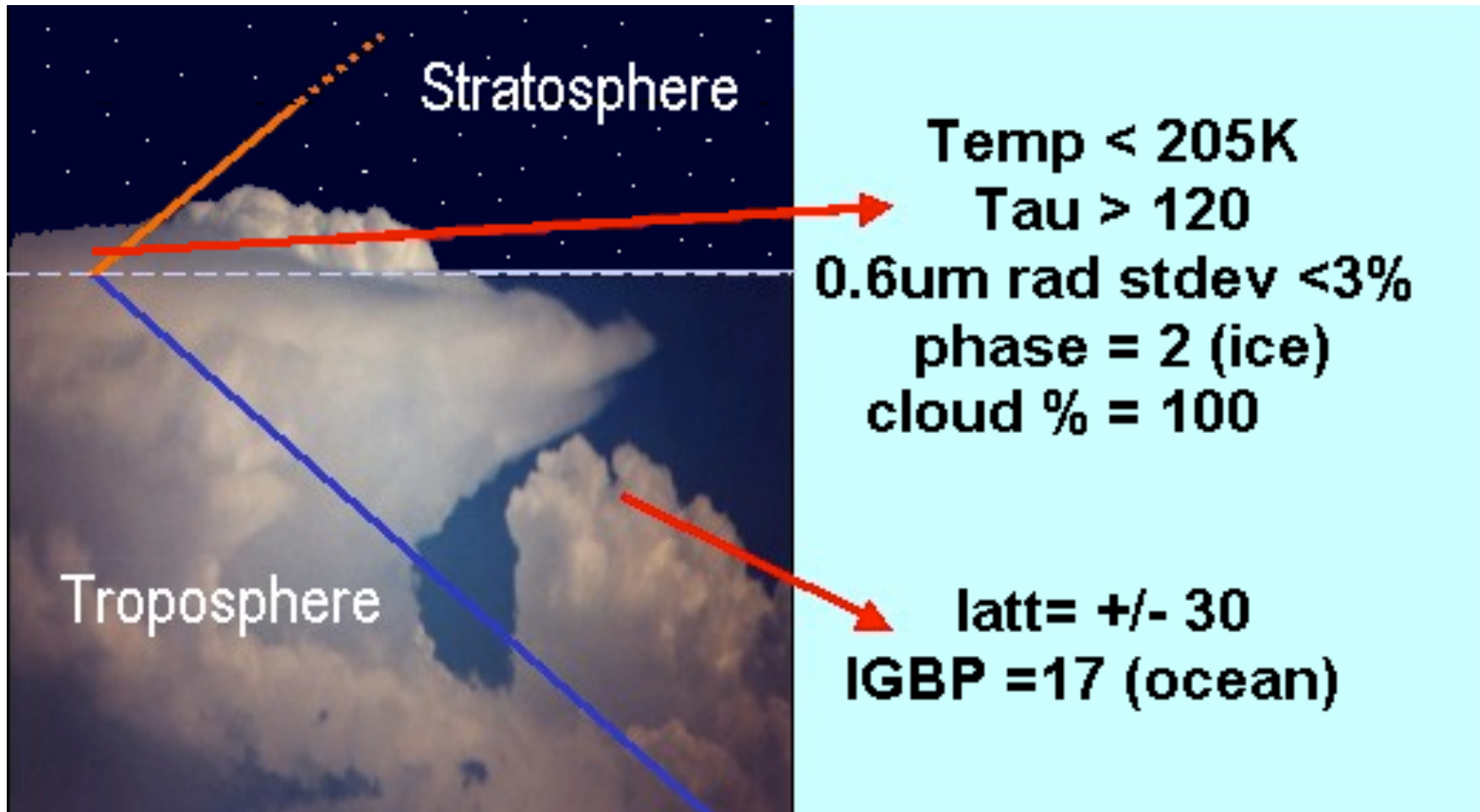
2. Edition 2 Un-filtered Direct Compare



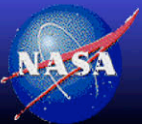
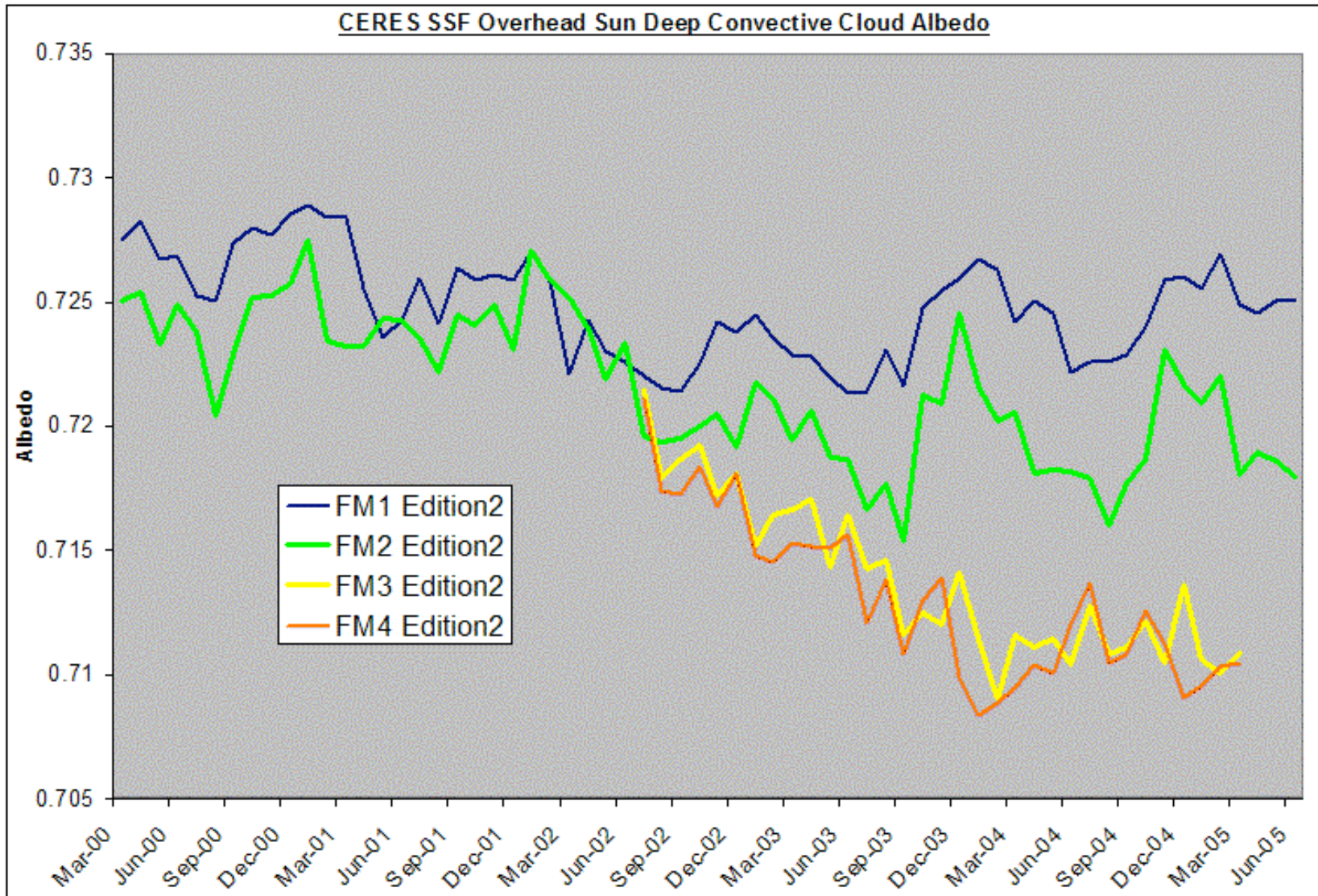
NASA Langley Research Center / Science Directorate



Use MODIS to find SW footprints of thickest, coldest and most uniform tropical ocean Deep Convective Clouds



Edition 2 DCC albedo is a metric for instrument degradation



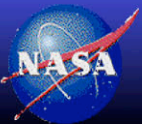
NASA Langley Research Center / Science Directorate



1 RAPS, 1 Xtrack instrument at anytime. Direct Compare (DC) is the ratio of simultaneous nadir un-filtered radiance



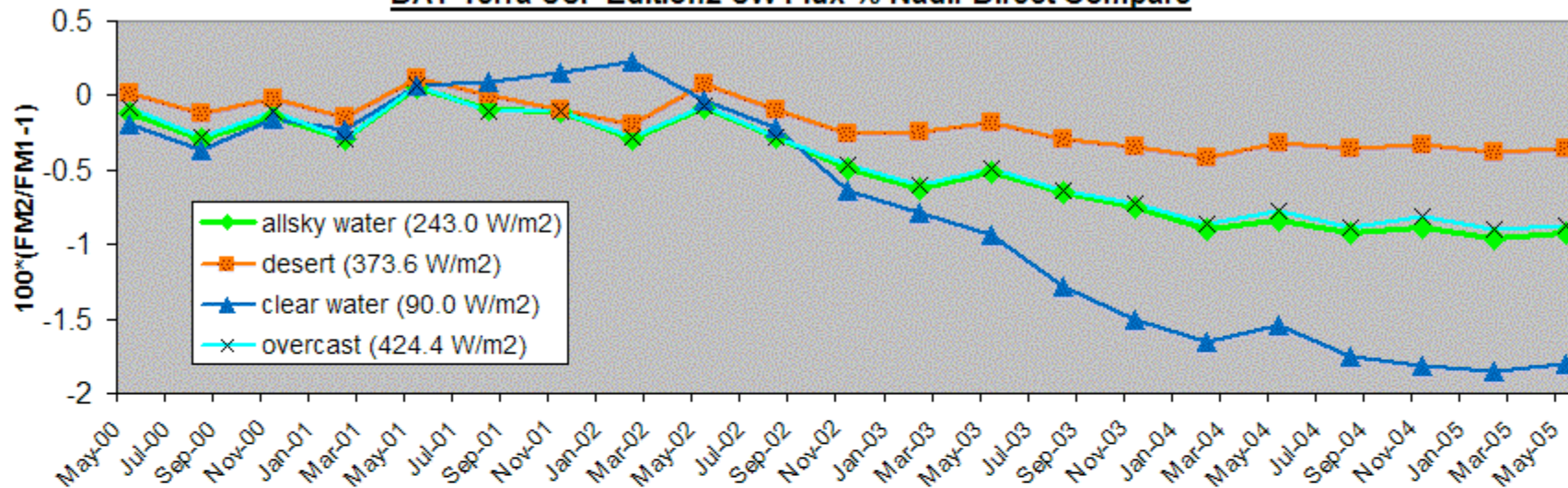
Instrument diagnostic that is **INDEPENDENT OF CLIMATE**



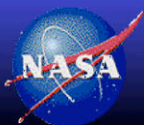
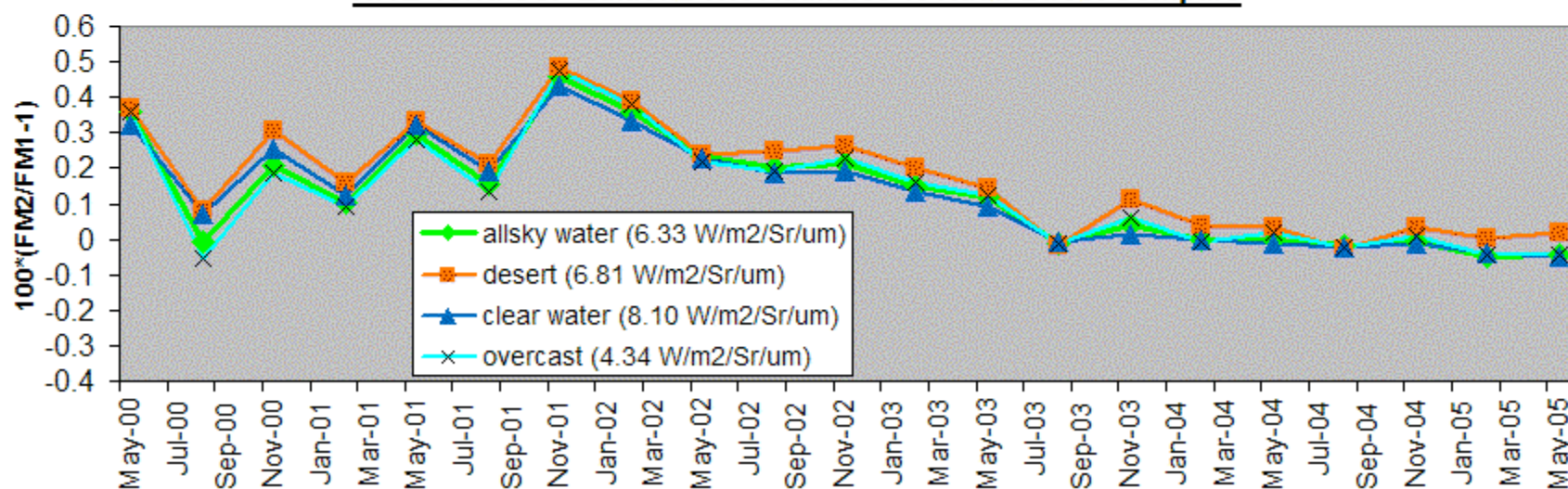
NASA Langley Research Center / Science Directorate



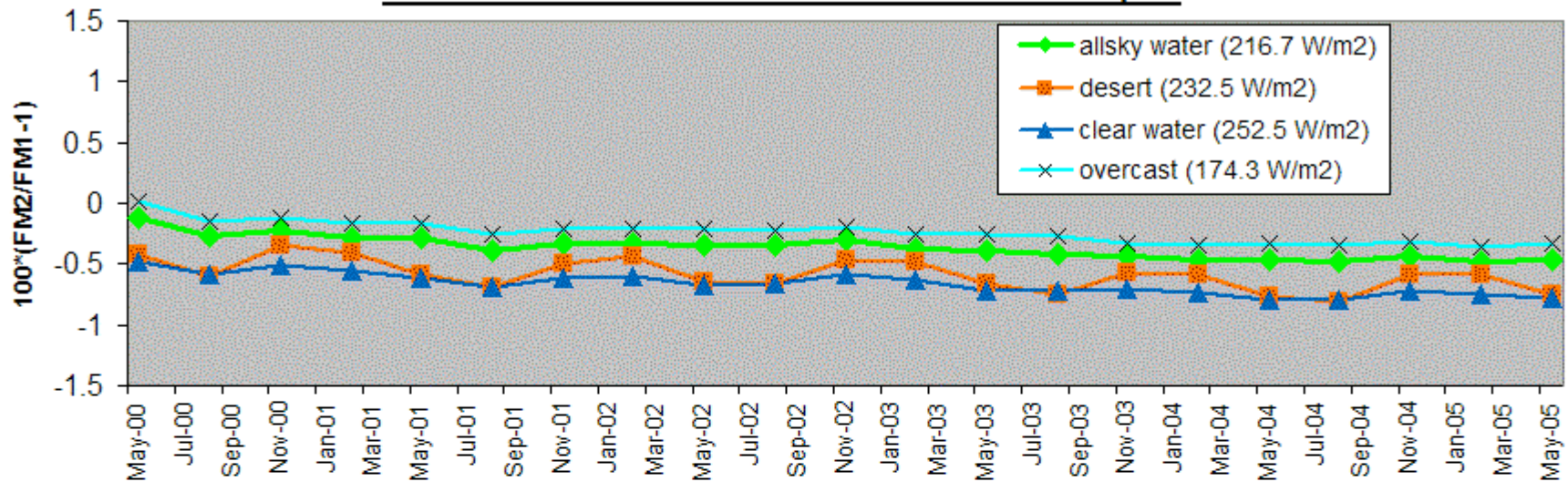
DAY Terra SSF Edition2 SW Flux % Nadir Direct Compare



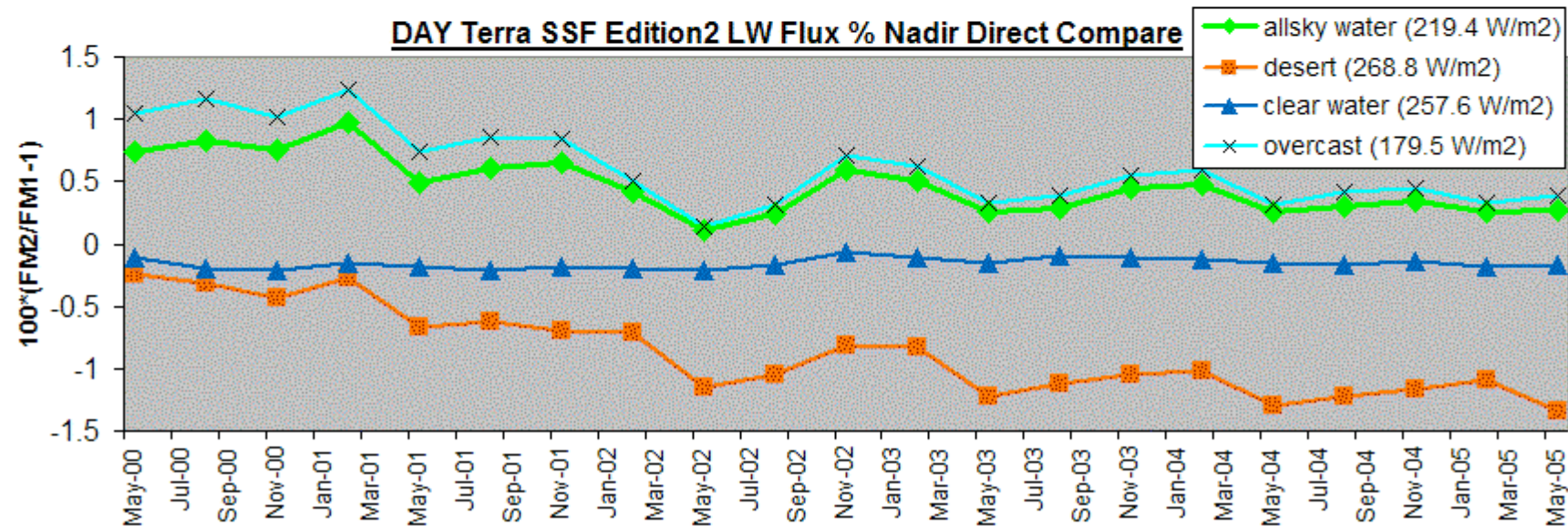
NITE Terra SSF Edition2 WN Radiance % Nadir Direct Compare



NITE Terra SSF Edition2 LW Flux % Nadir Direct Compare

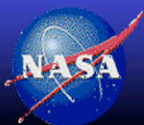
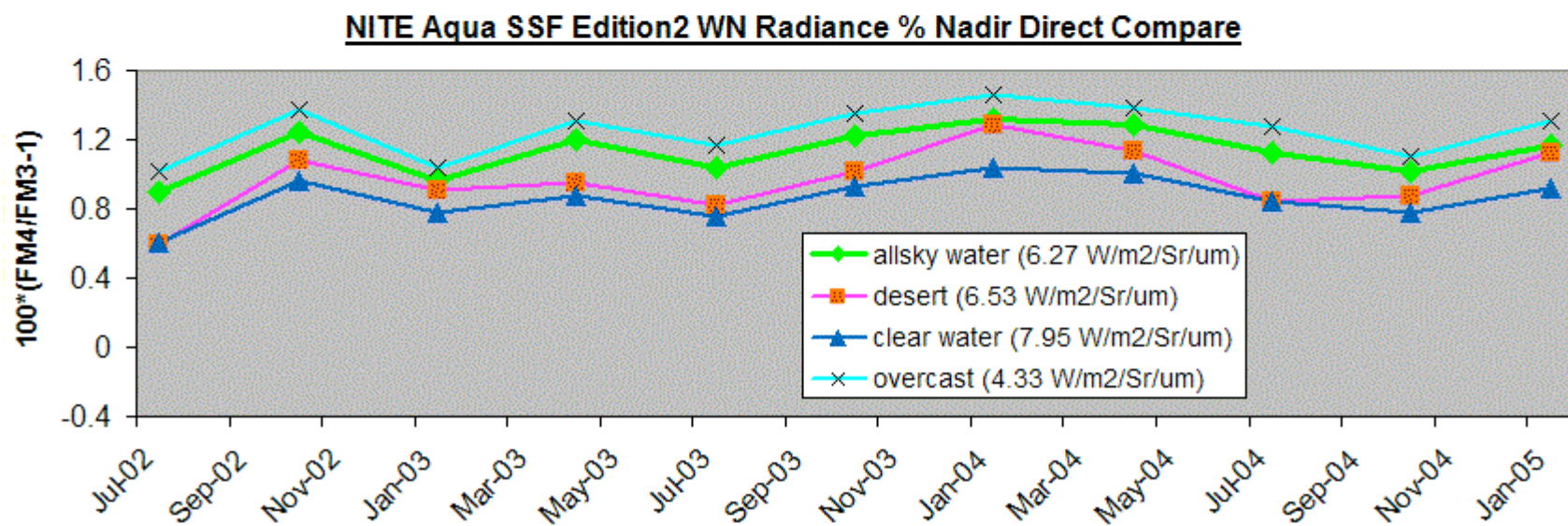
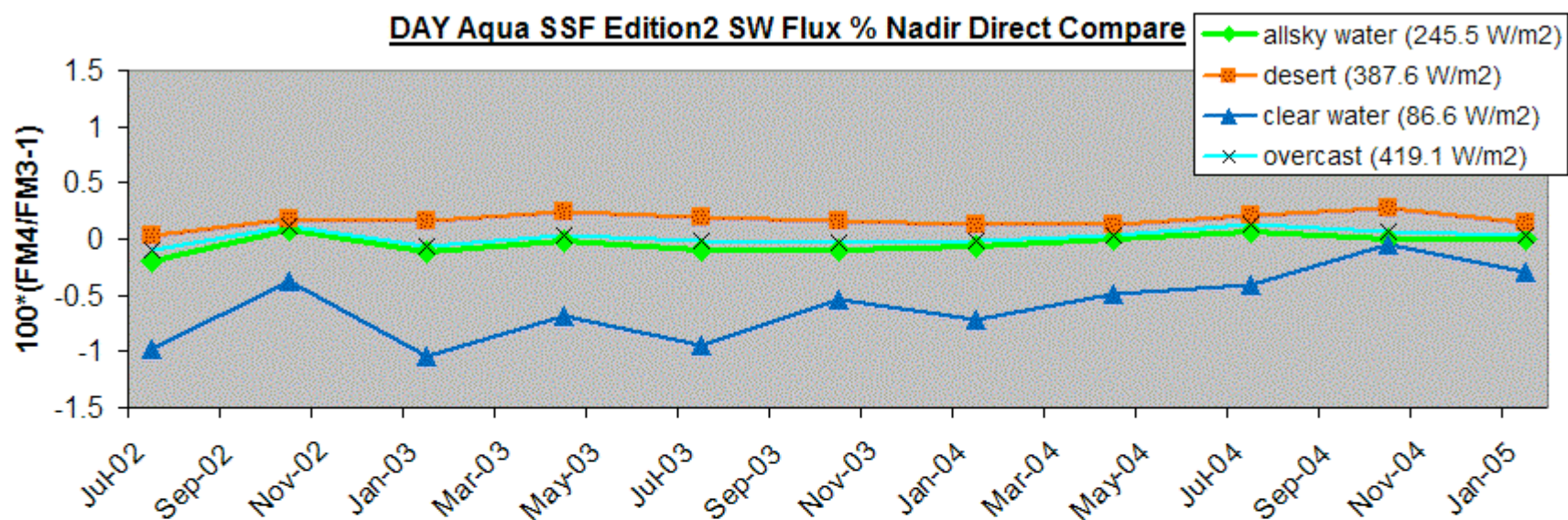


DAY Terra SSF Edition2 LW Flux % Nadir Direct Compare



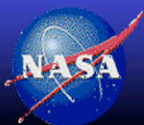
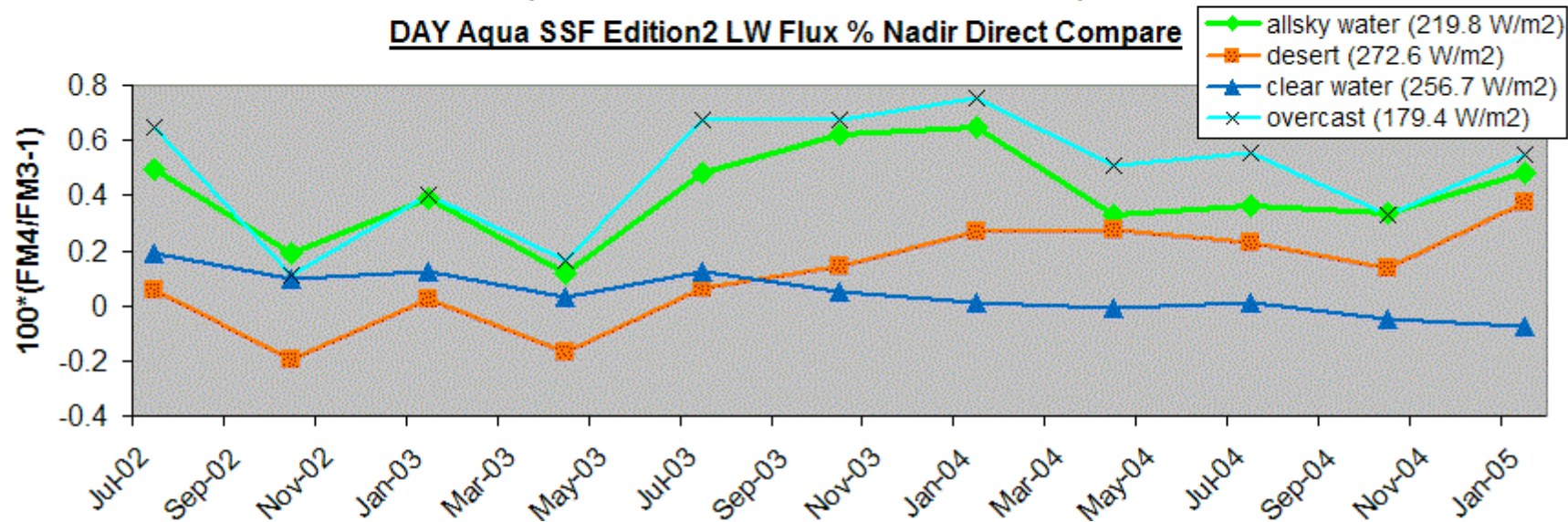
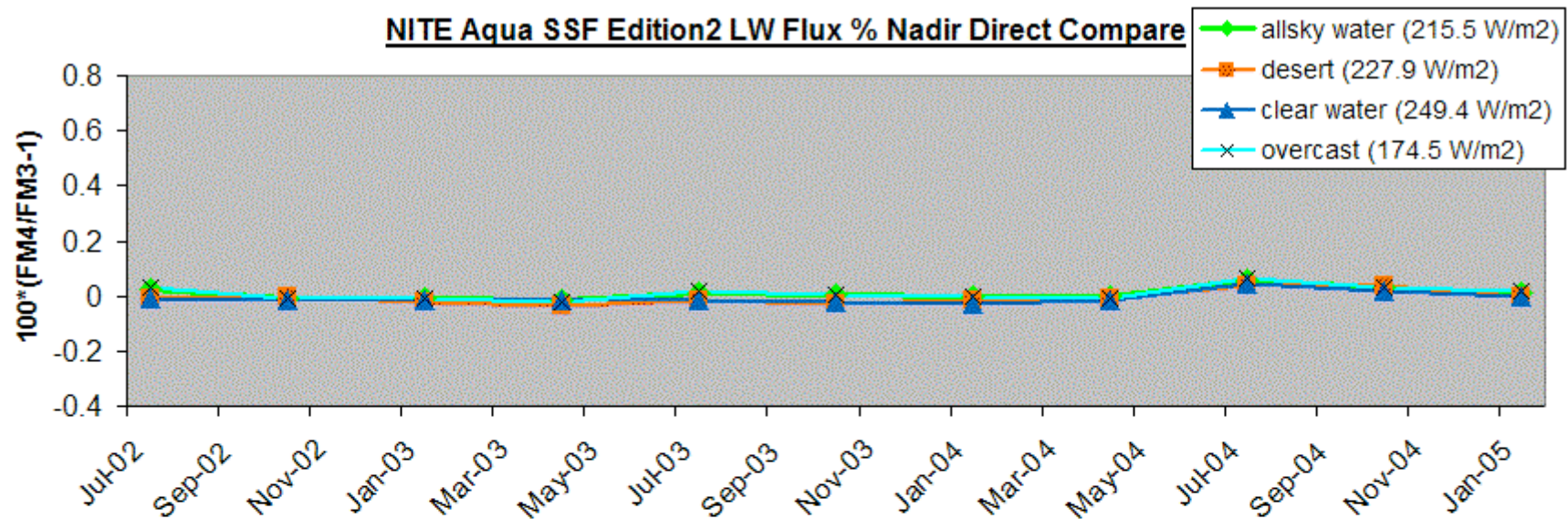
NASA Langley Research Center / Science Directorate





NASA Langley Research Center / Science Directorate



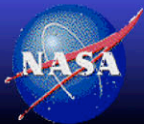


NASA Langley Research Center / Science Directorate



SW Edition 3 calibration

Model Derived Gains and Spectral response

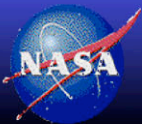


NASA Langley Research Center / Science Directorate



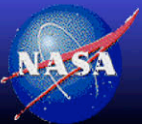
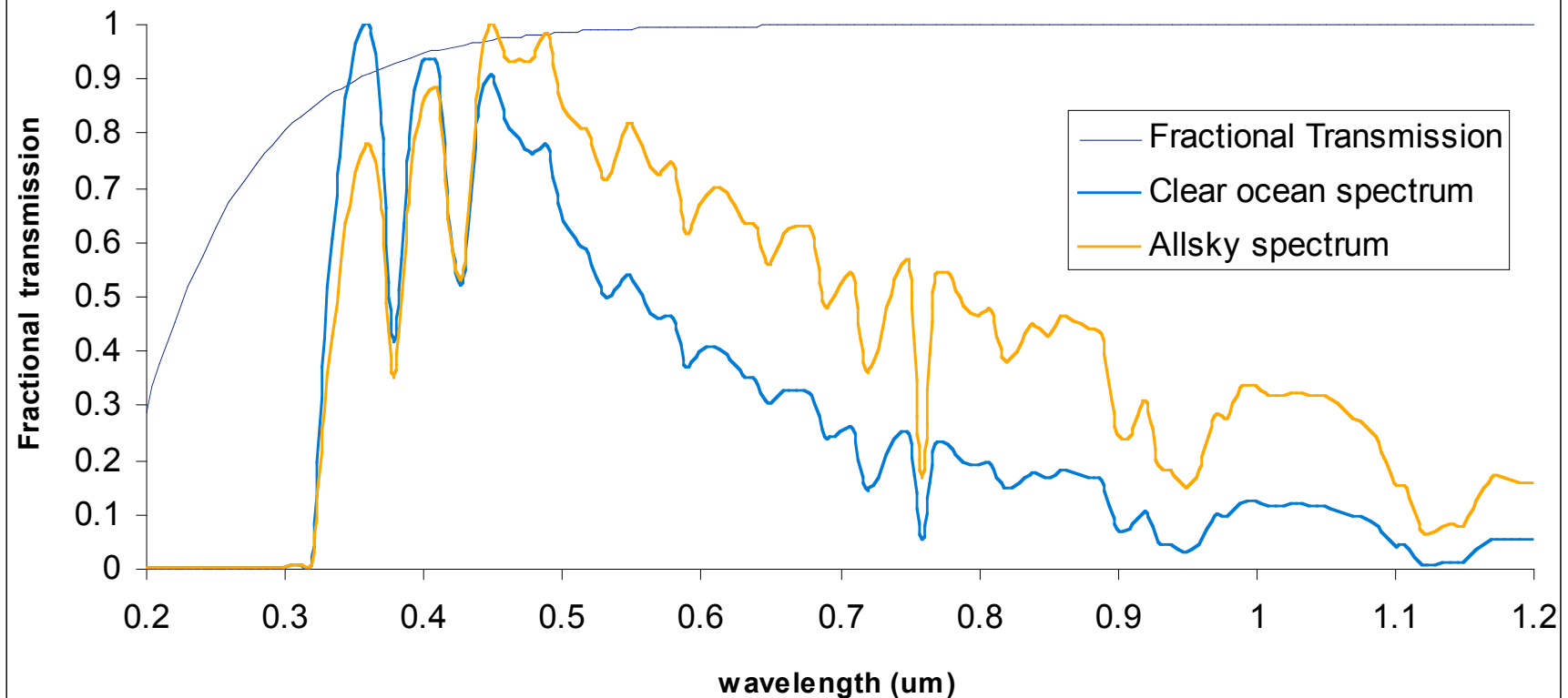
The deficiencies in the Aqua Rev1 adjustment suggest significant cross-track instrument darkening or on-board lamp drift ($< 1\%$). Hence for Edition 3, a new calibration methodology is needed:

- **Deep Convective Cloud albedo:** Find the albedo of the coldest, brightest and most uniform clouds in the tropics. These become your ‘solar diffuser’ stability metric to replace MAM data.
- **Constrain contamination model using direct compare:** Contamination/UV exposure model is adjusted to best match the direct compare of nadir footprints between two instruments on the same platform. Use of clear ocean and allsky scenes allows the model to determine coloration of spectral response changes.



$$D(\lambda) = [1 - M.e^{-\alpha\lambda}]$$

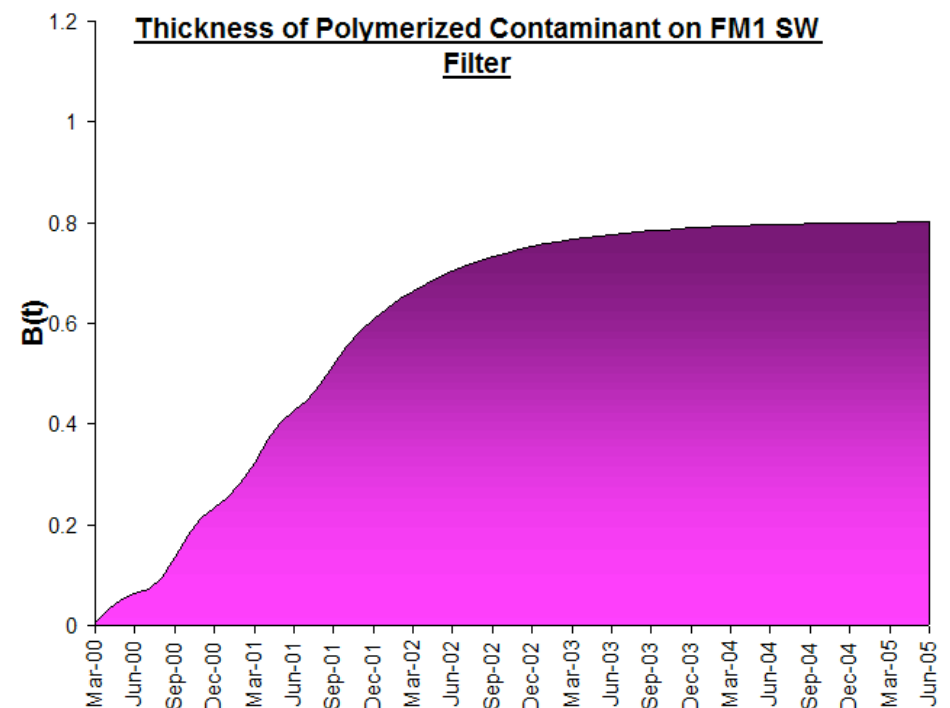
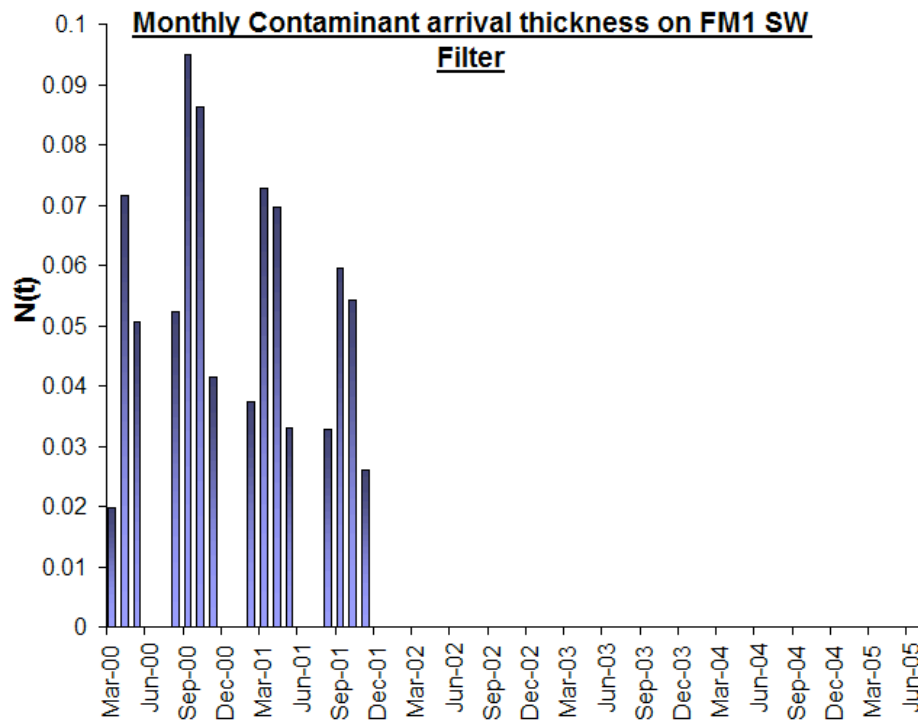
Plot of normalized Earth spectra compared to fractional transmission of polymerized contaminant transmission (1um thick)



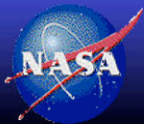
NASA Langley Research Center / Science Directorate



FM1 Contaminant Thickness



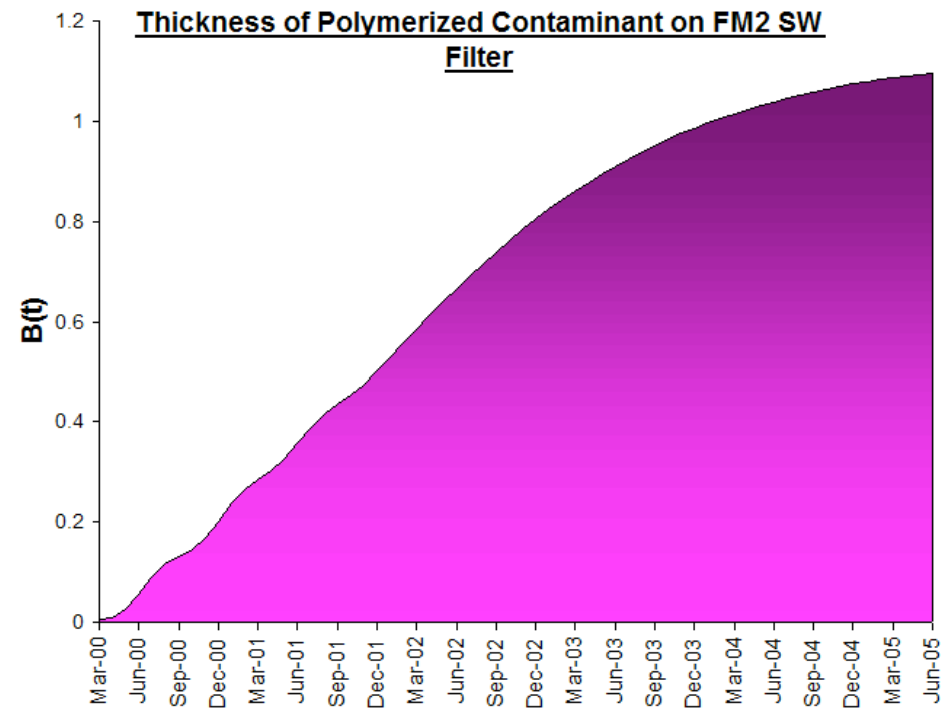
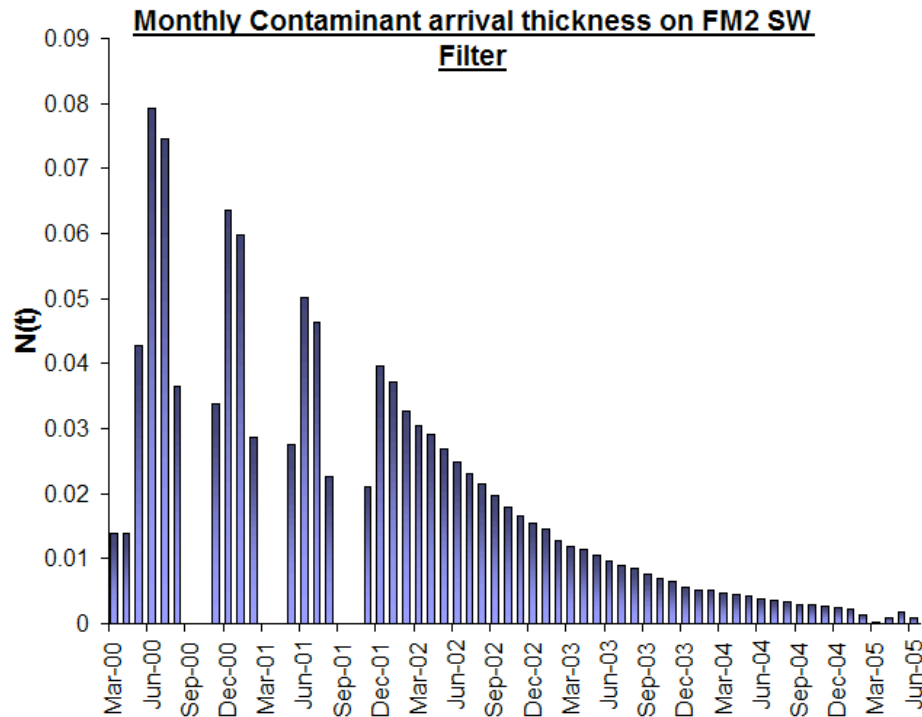
$$\frac{dB(t)}{dt} = \rho \left[\int_0^t N(\xi) d\xi - B(t) \right] + \beta \cdot N(t)$$



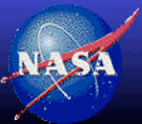
NASA Langley Research Center / Science Directorate



FM2 Contaminant Thickness



$$\frac{dB(t)}{dt} = \rho \left[\int_0^t N(\xi) d\xi - B(t) \right] + \beta \cdot N(t)$$



NASA Langley Research Center / Science Directorate



Model Spectral Darkening found from contaminant unit transmission and thickness $B(t)$:

$$D(\lambda, t) = [1 - M.e^{-\alpha\lambda}]^{B(t)}$$

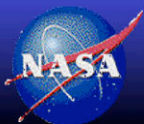
$$S^{ed3}(\lambda, t) = D(\lambda, t) \times S^{Gnd}(\lambda)$$

Detector Output = Gain * Filtered Radiance

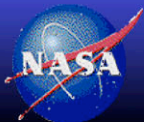
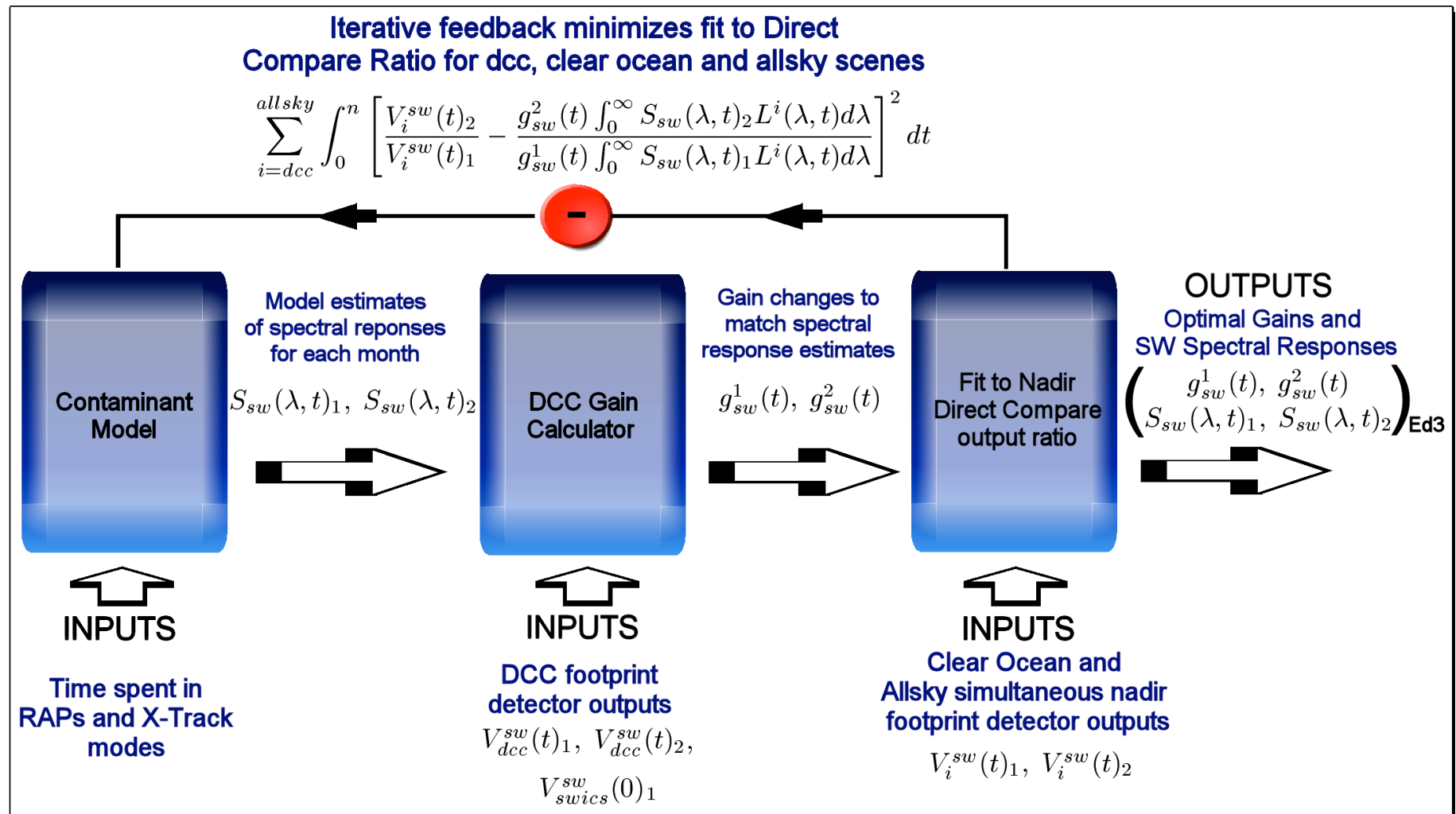
$$V = g \times \int_0^{\infty} S(\lambda)L(\lambda)d\lambda$$

Hence SW gain changes can be found using the DCC detector signal and modeled filtered radiance:

$$g_{sw}(t) = \frac{V_{dcc}^{sw}(t)}{\int_0^{\infty} S_{sw}^{ed3}(\lambda, t)L^{dcc}(\lambda, t)d\lambda}$$

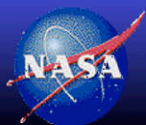
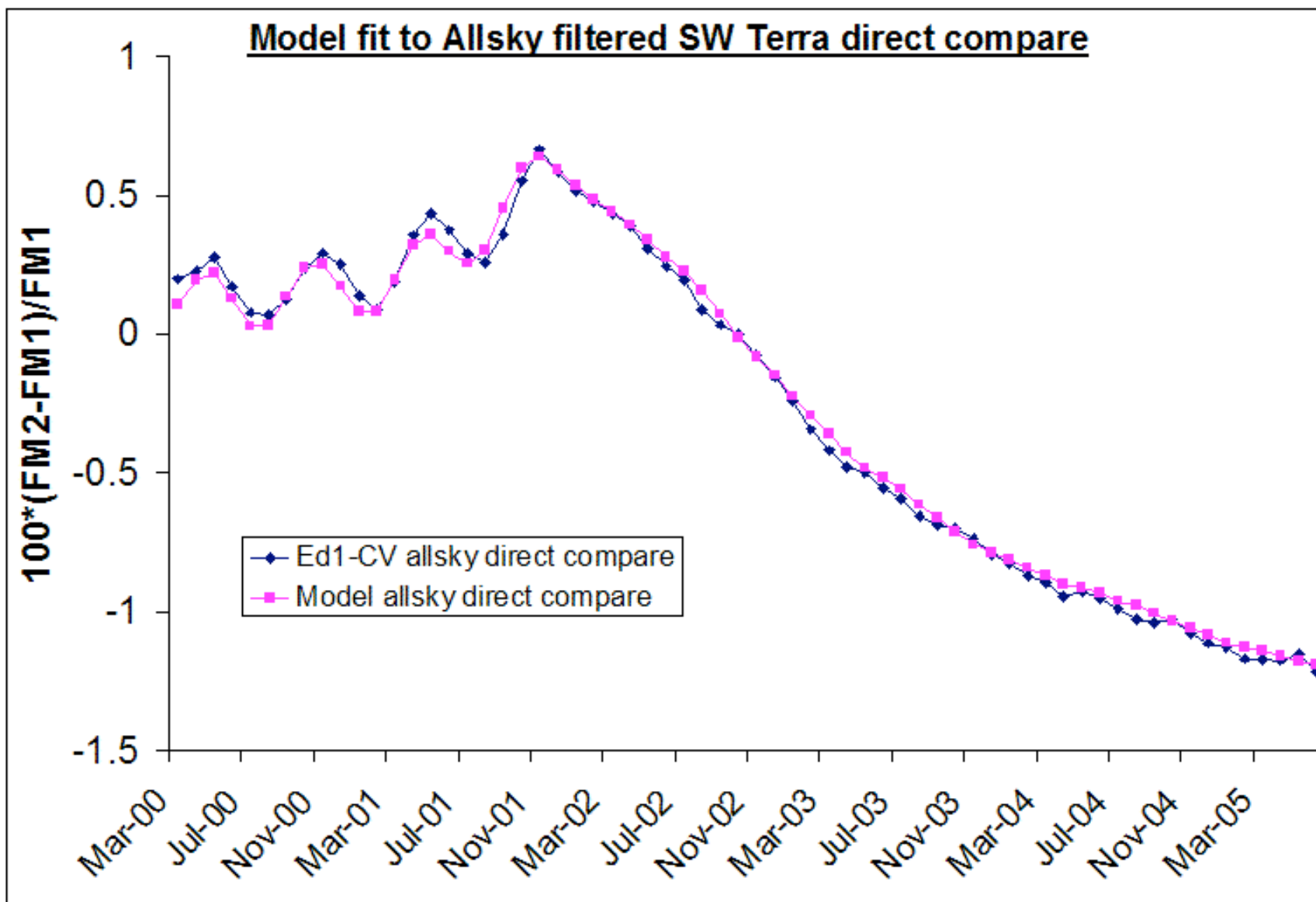


Run iterative model for all 'n' months available:



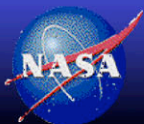
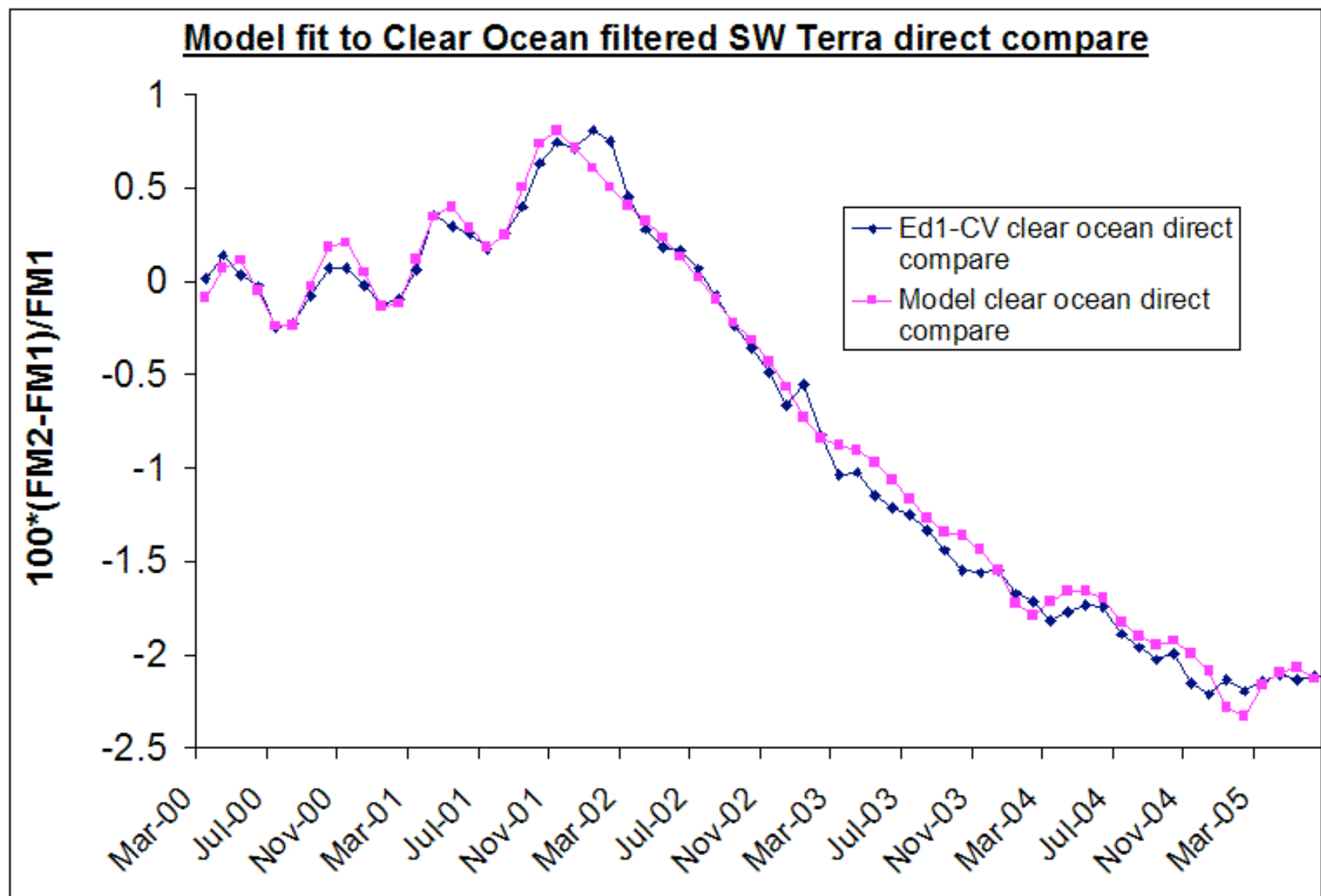
NASA Langley Research Center / Science Directorate





NASA Langley Research Center / Science Directorate

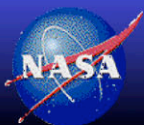
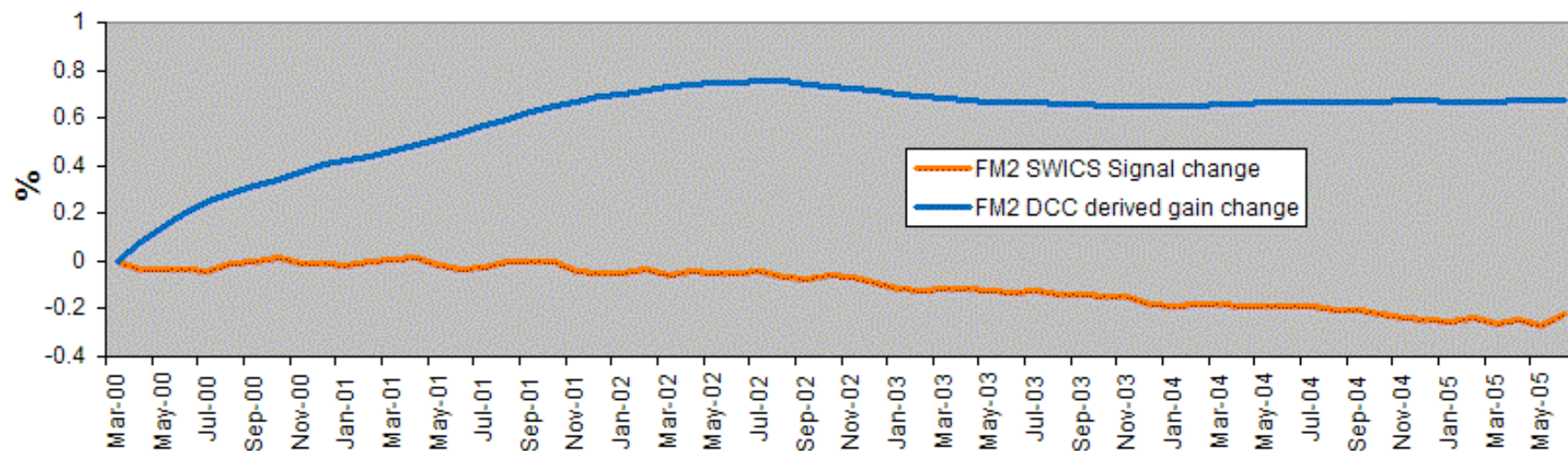
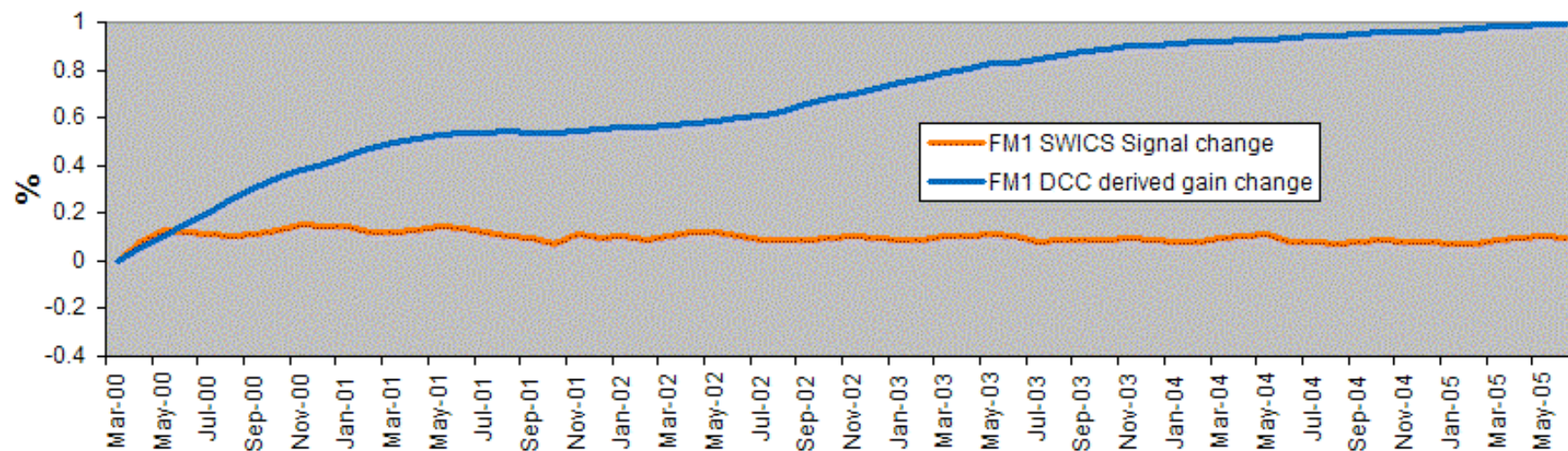




NASA Langley Research Center / Science Directorate



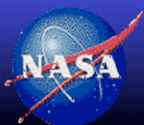
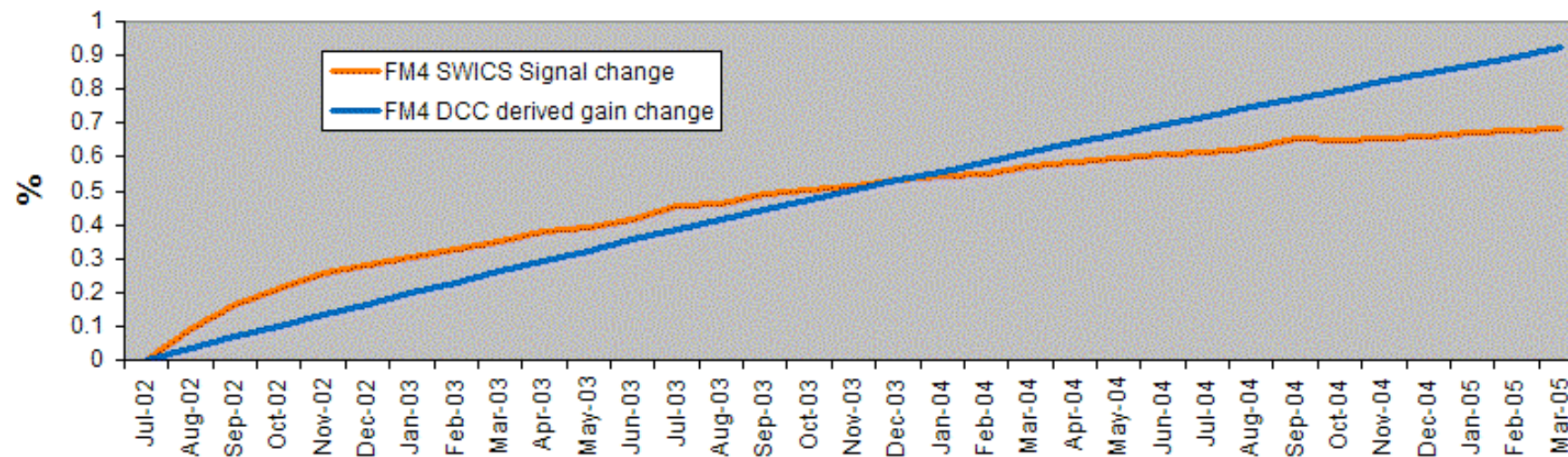
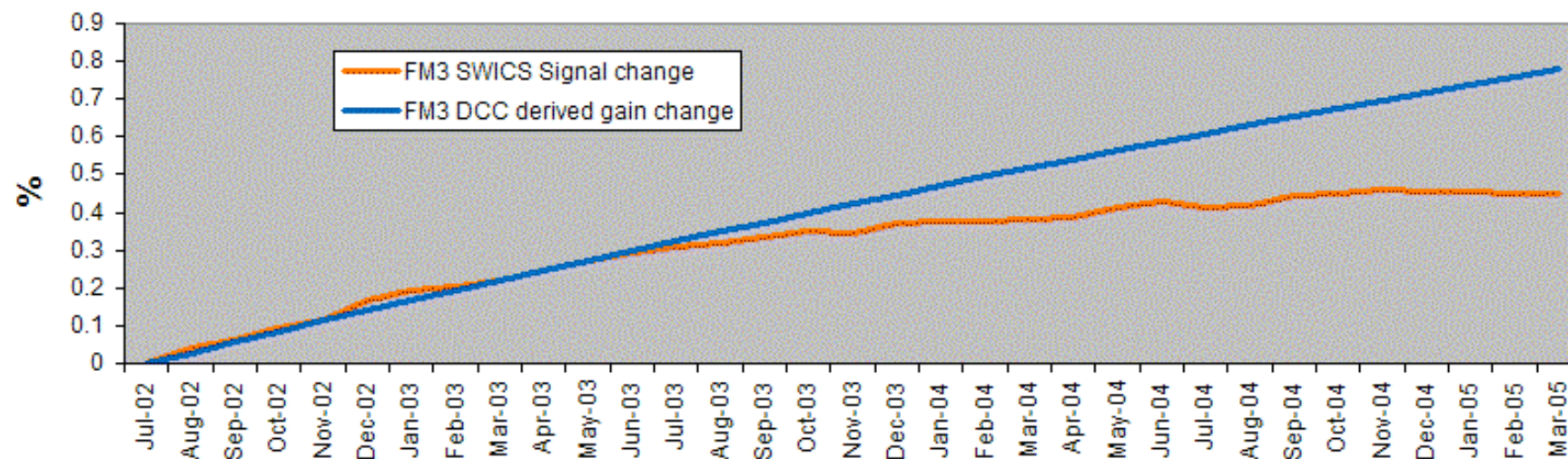
Terra Deep Convective Cloud SW % Gain change from Mission start



NASA Langley Research Center / Science Directorate



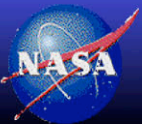
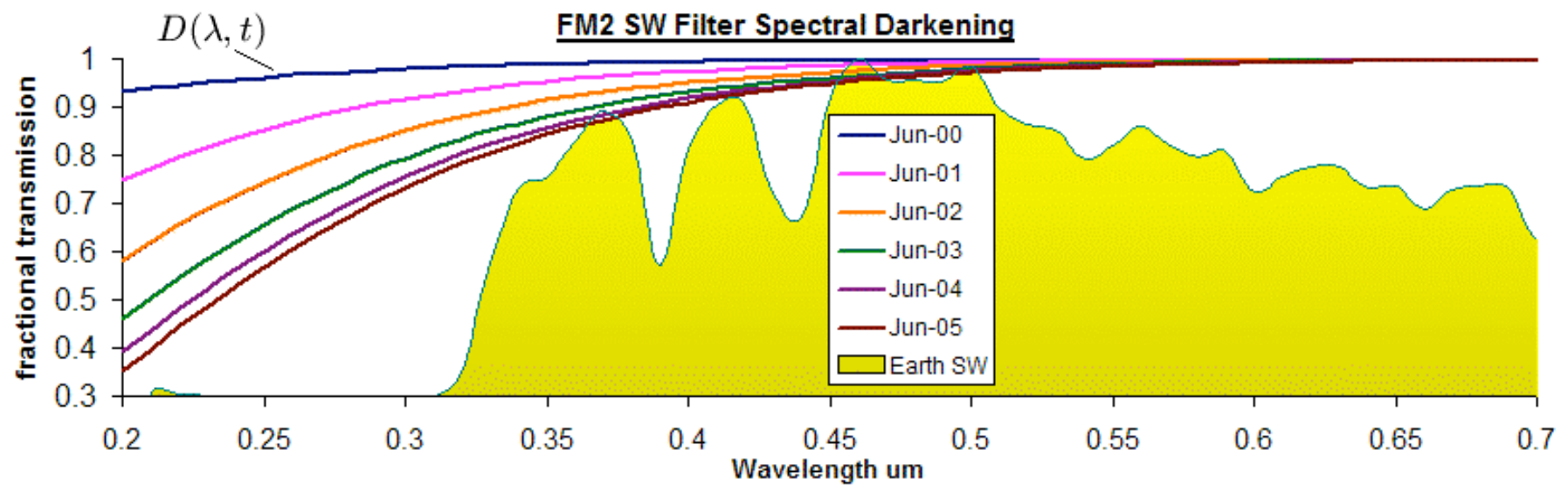
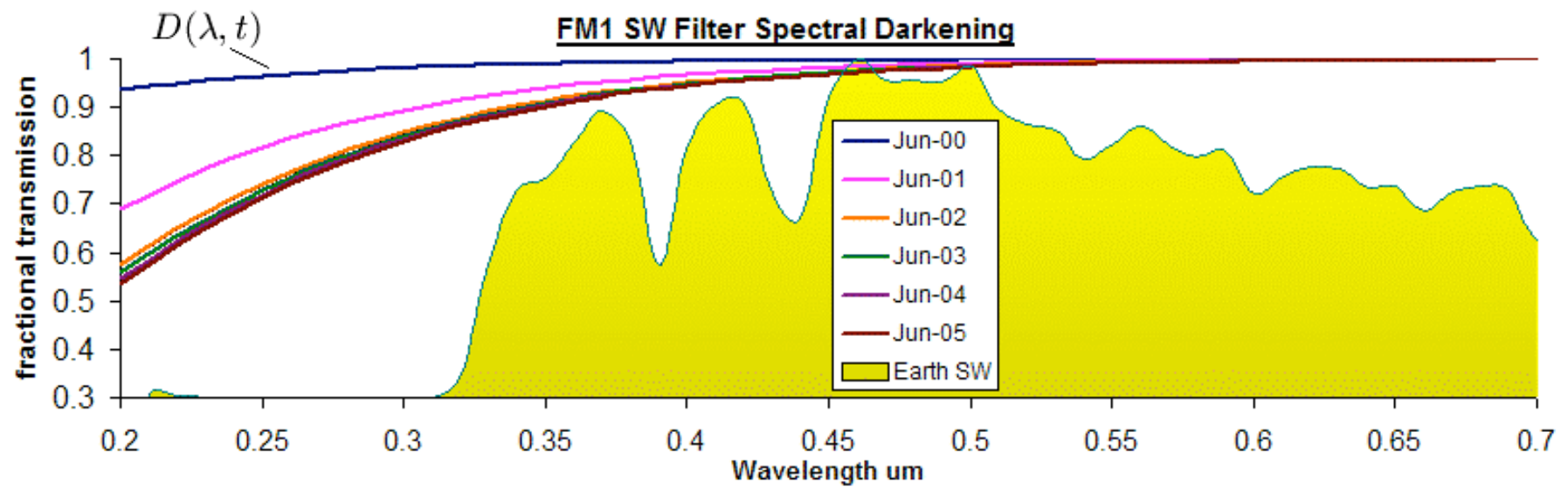
Aqua Deep Convective Cloud SW % Gain change from Mission start



NASA Langley Research Center / Science Directorate



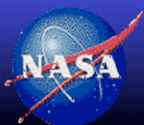
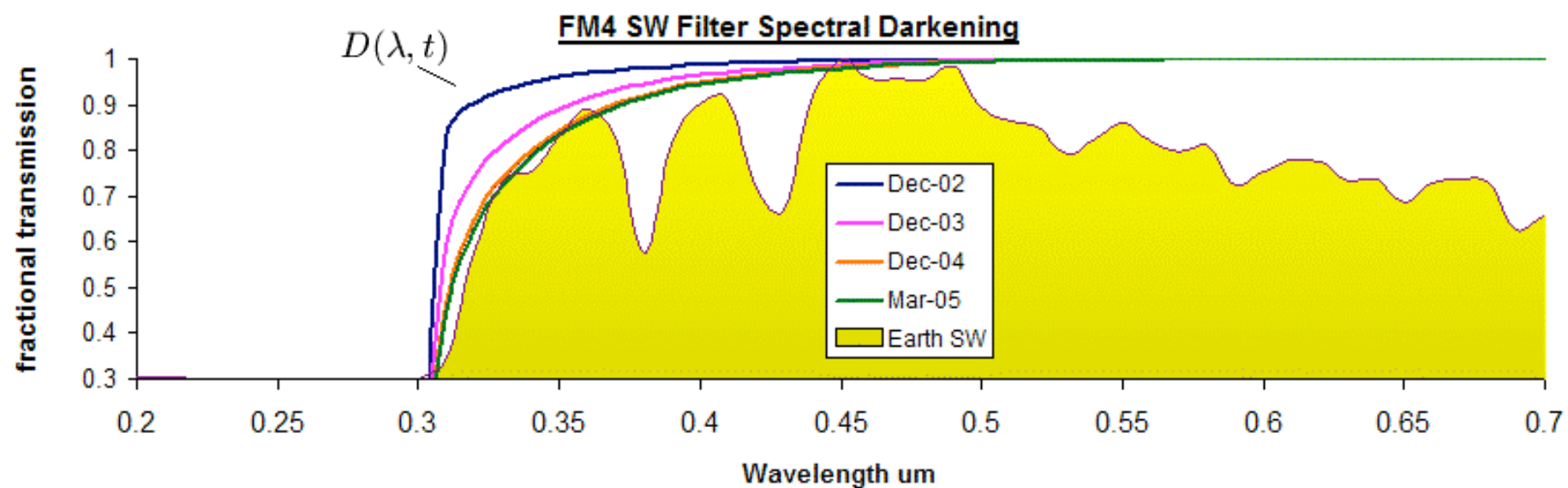
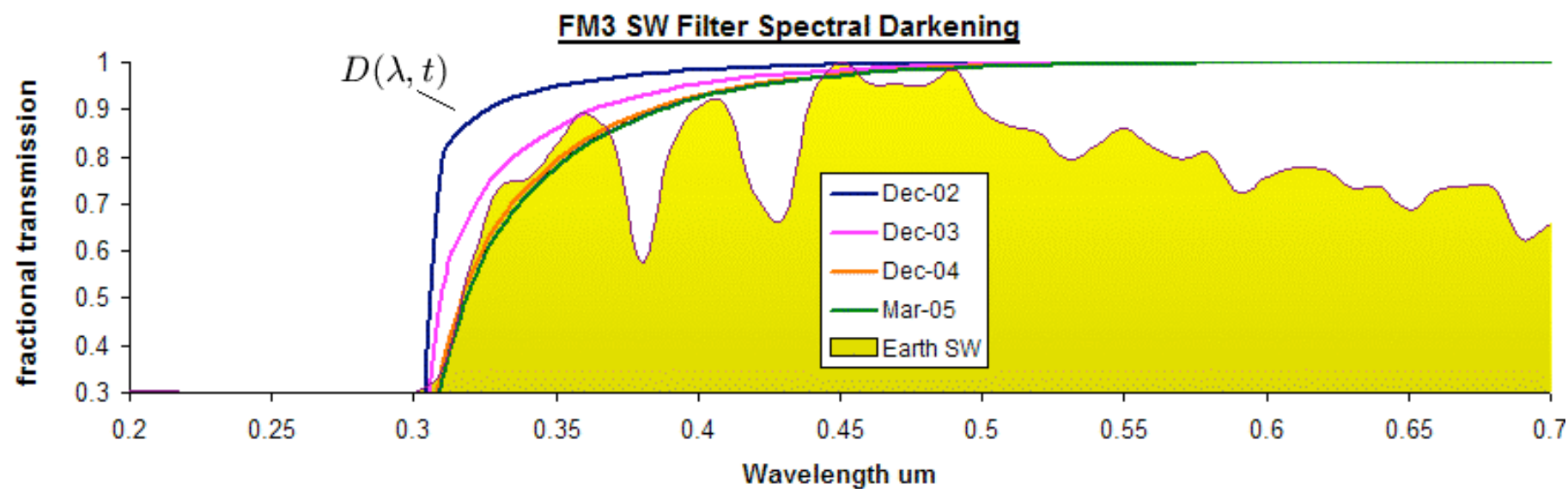
TERRA



NASA Langley Research Center / Science Directorate



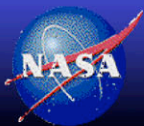
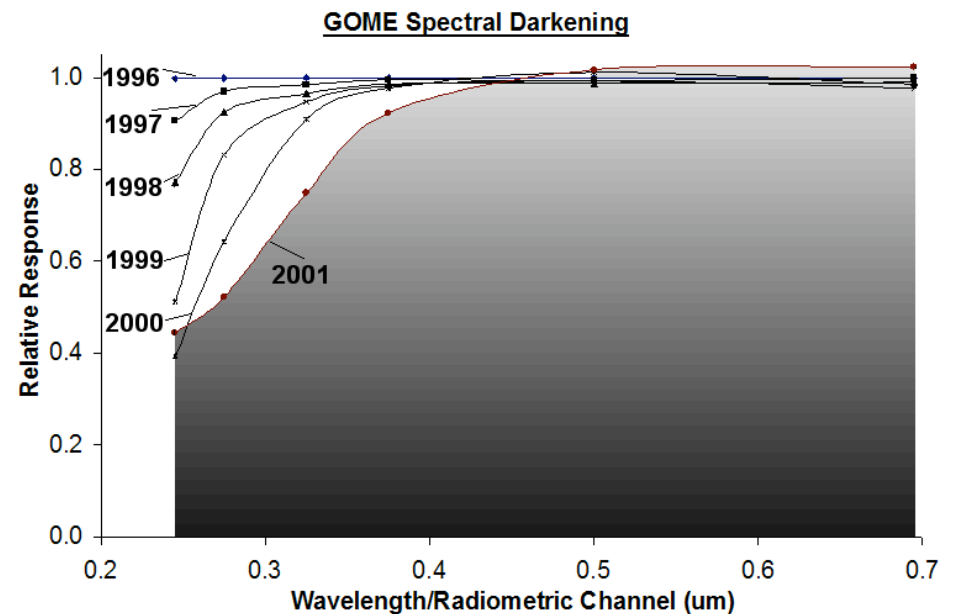
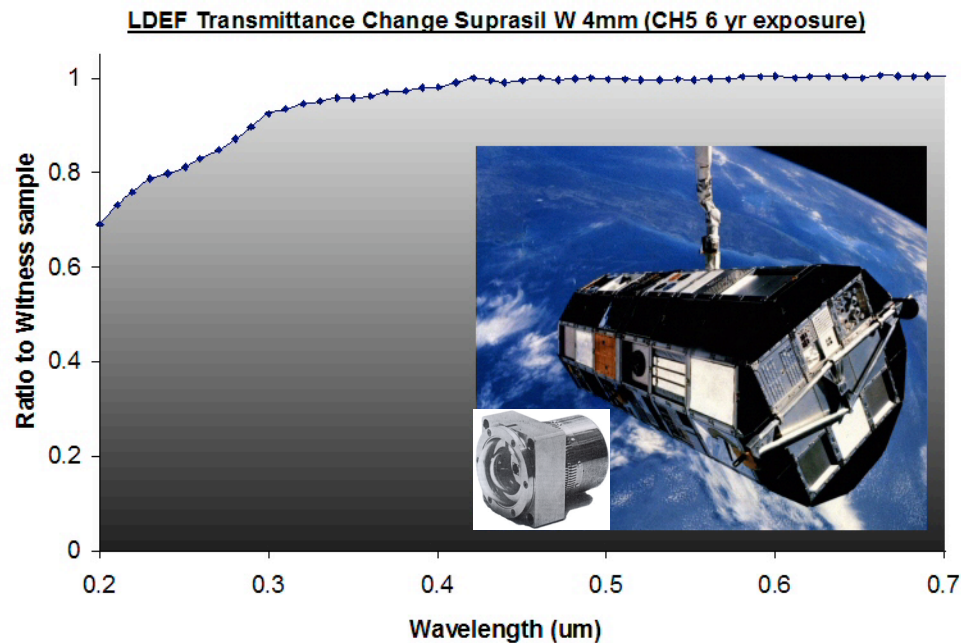
AQUA



NASA Langley Research Center / Science Directorate



This matches the shape of spectral darkening occurring on LDEF and GOME:

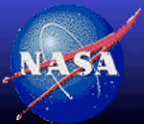


NASA Langley Research Center / Science Directorate



Total and WN Edition 3 calibration

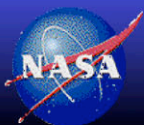
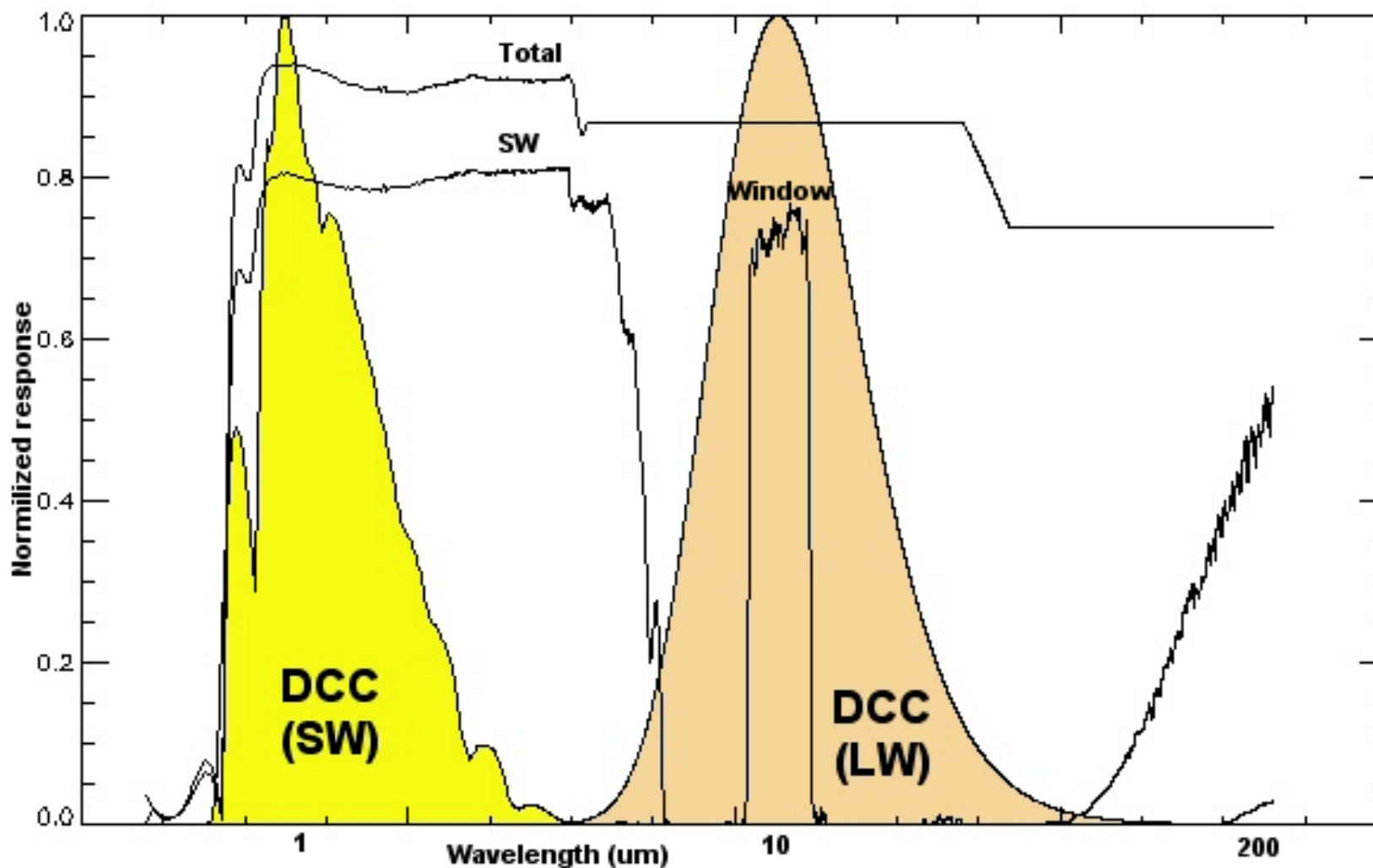
**ICM Derived Gains and Spectral response
based on SW spectral darkening**



NASA Langley Research Center / Science Directorate



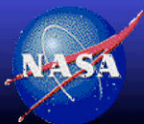
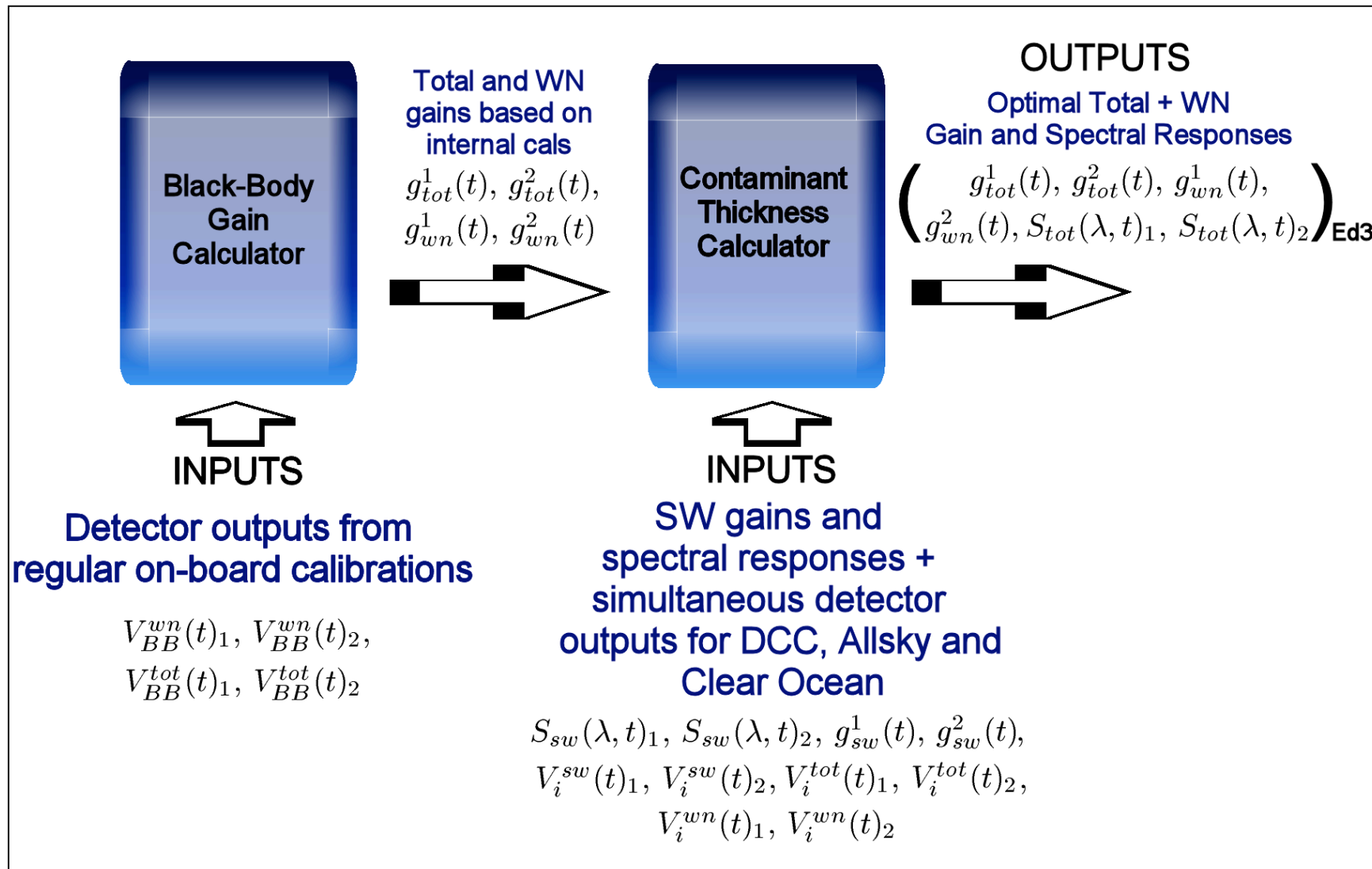
SW channel and SW portion of Total must balance exactly to give accurate daytime LW measurement:



NASA Langley Research Center / Science Directorate



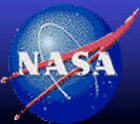
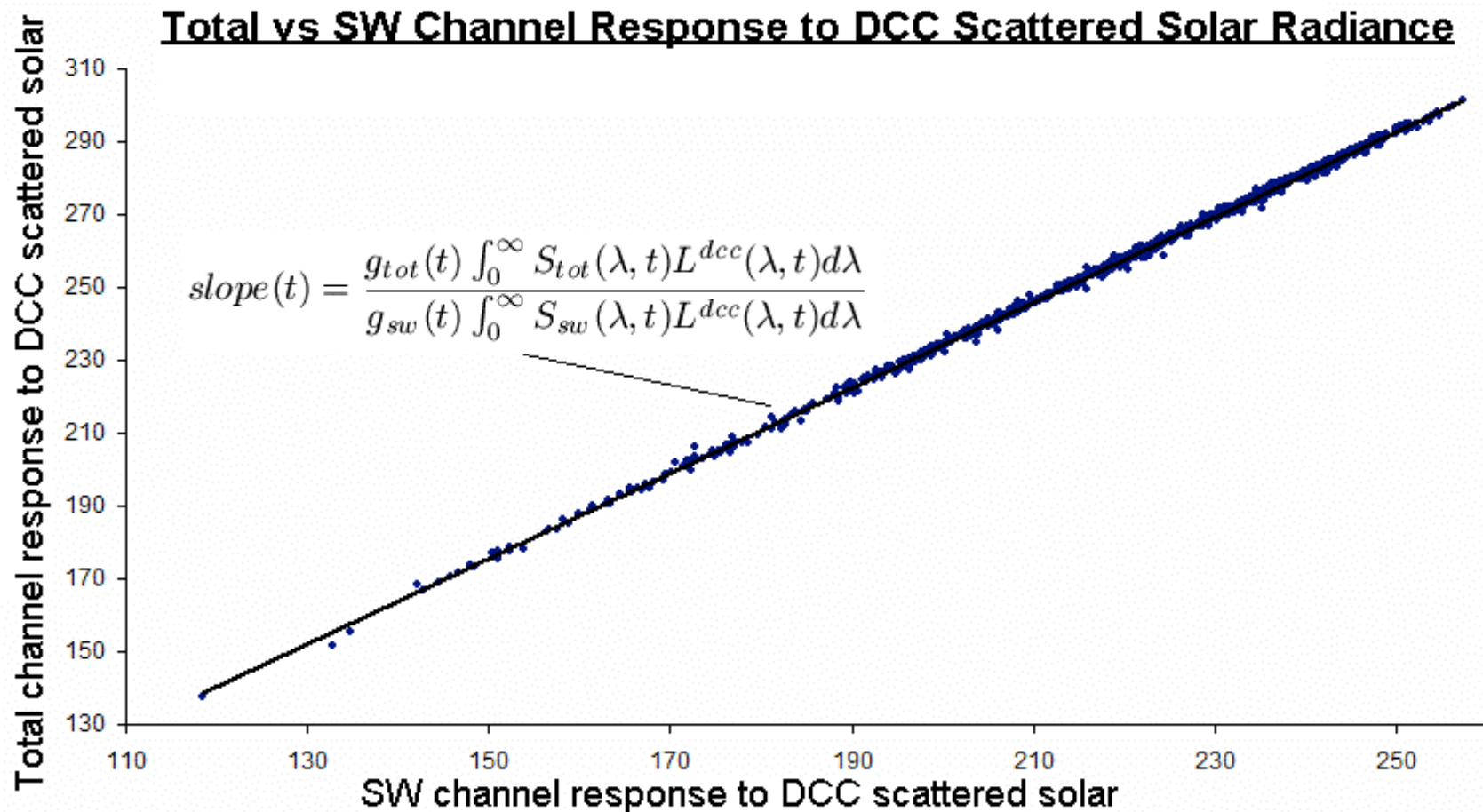
Use BB data to find gains then apply SW darkening to Total channel:



NASA Langley Research Center / Science Directorate



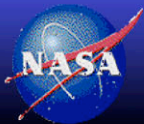
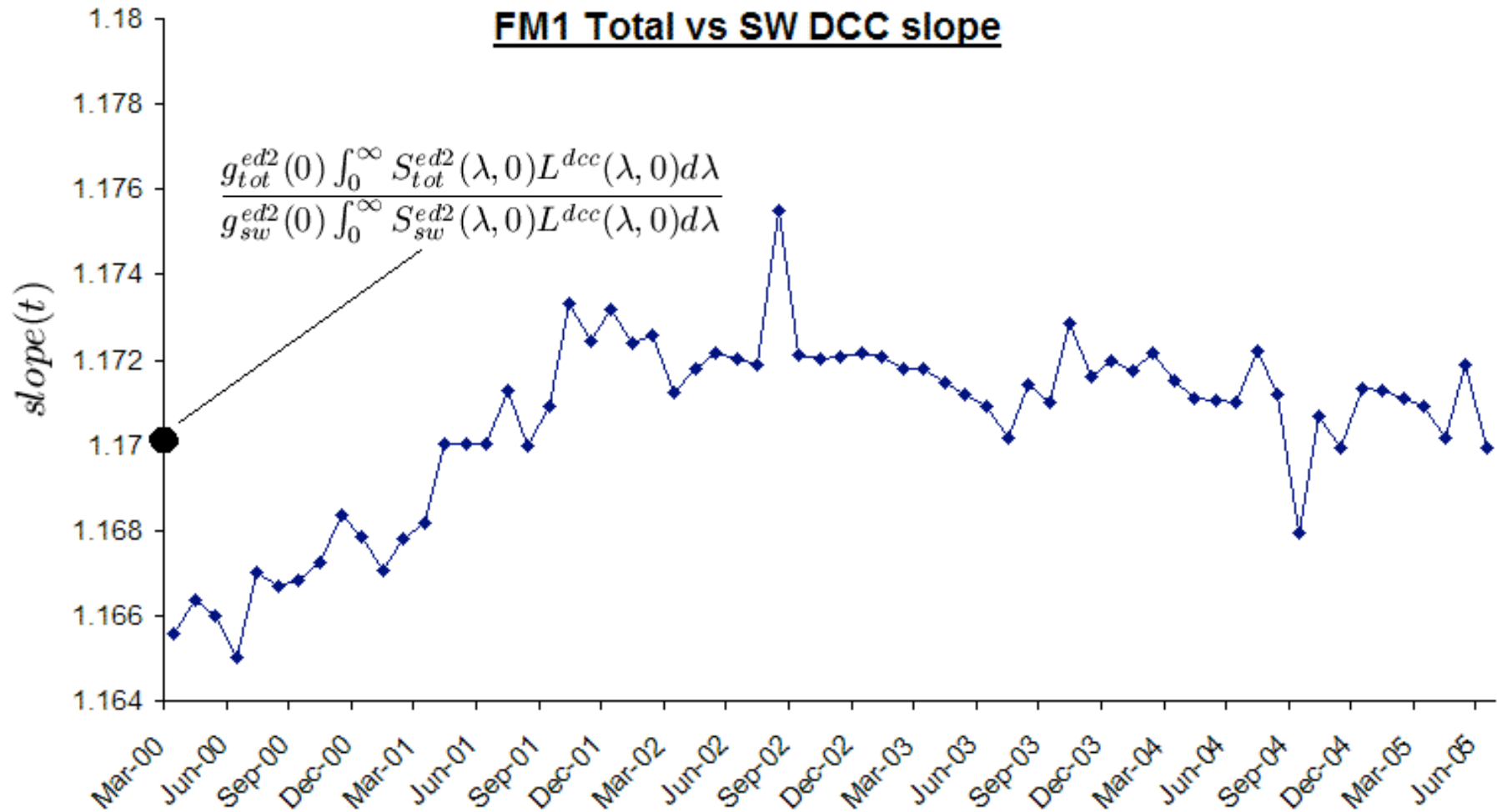
During day use window channel to remove the filtered LW DCC signal from filtered radiances. Compare Total and SW nadir response to DCC SW radiance:



NASA Langley Research Center / Science Directorate

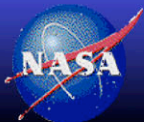
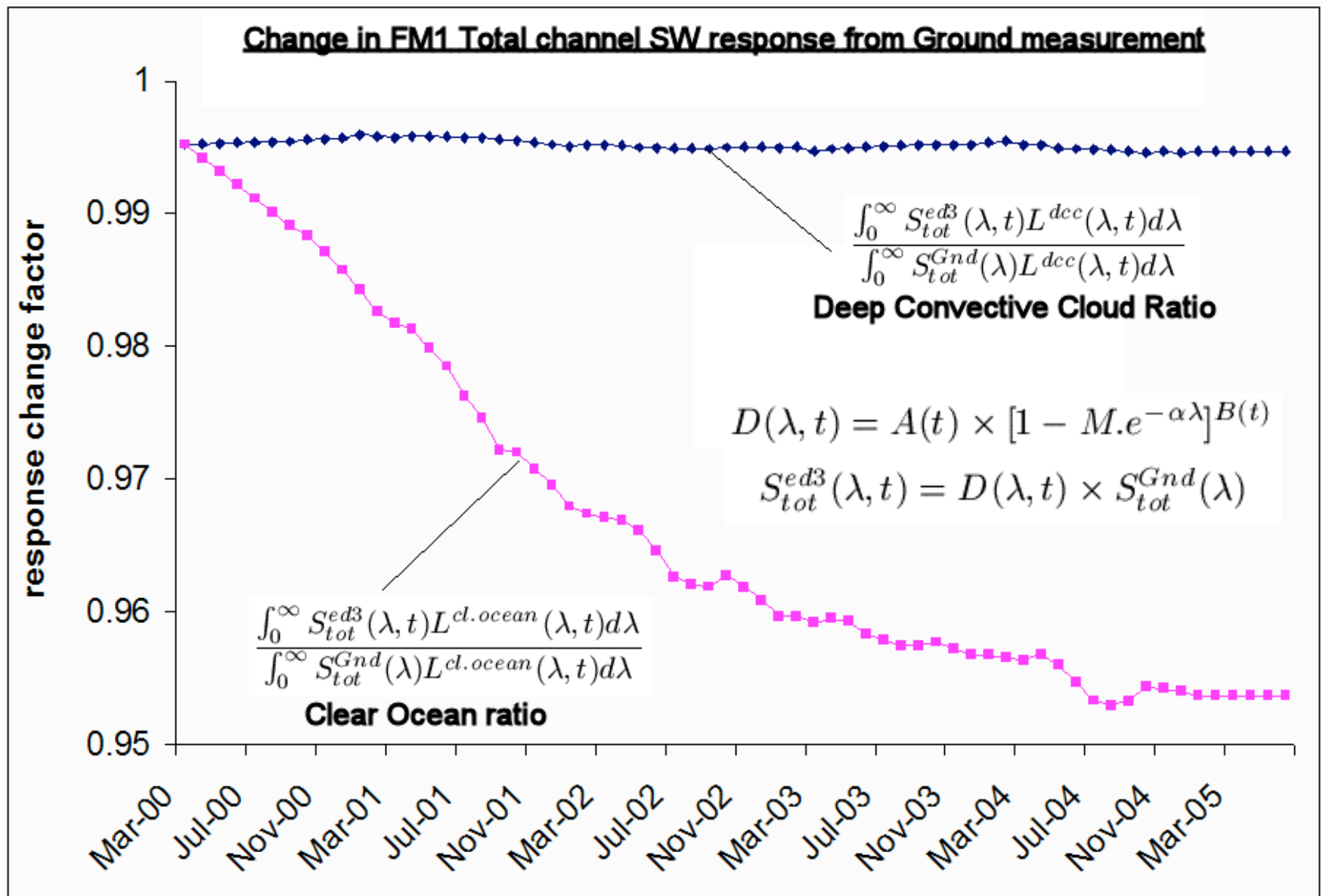


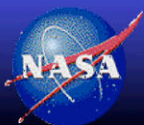
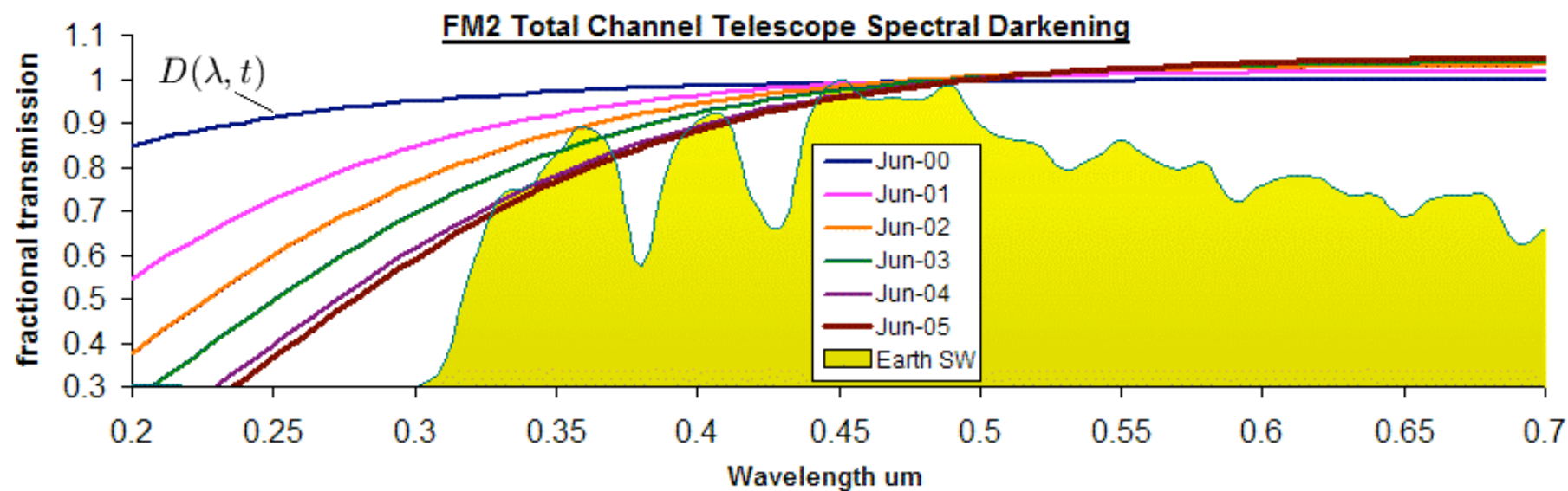
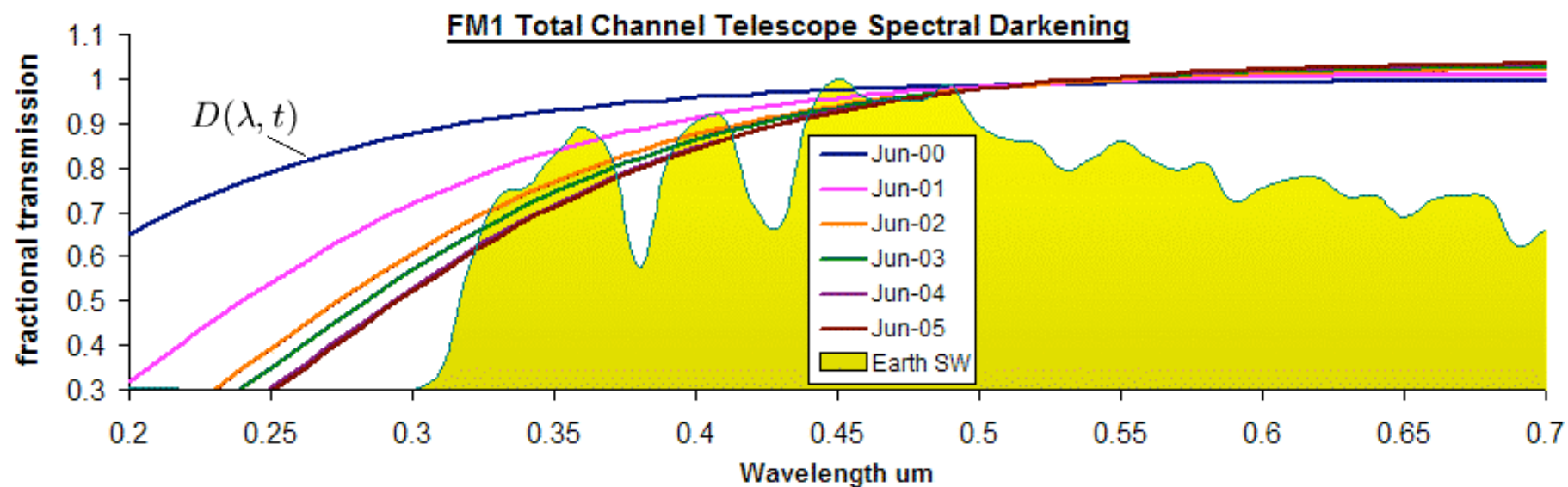
TERRA



NASA Langley Research Center / Science Directorate

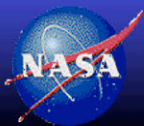
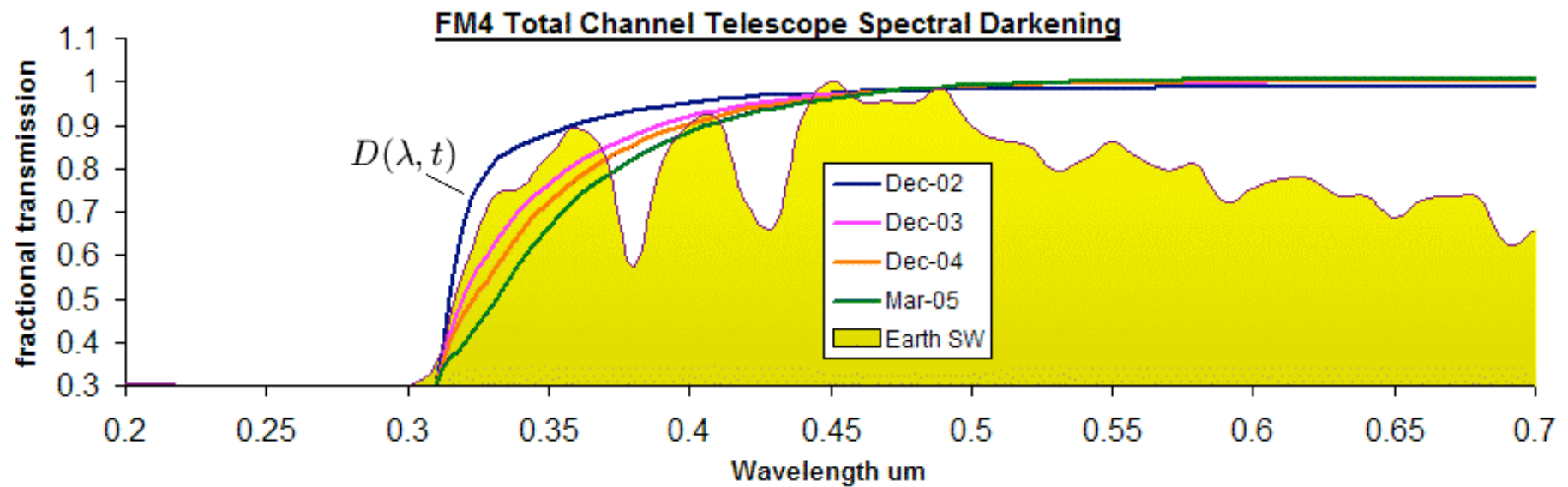
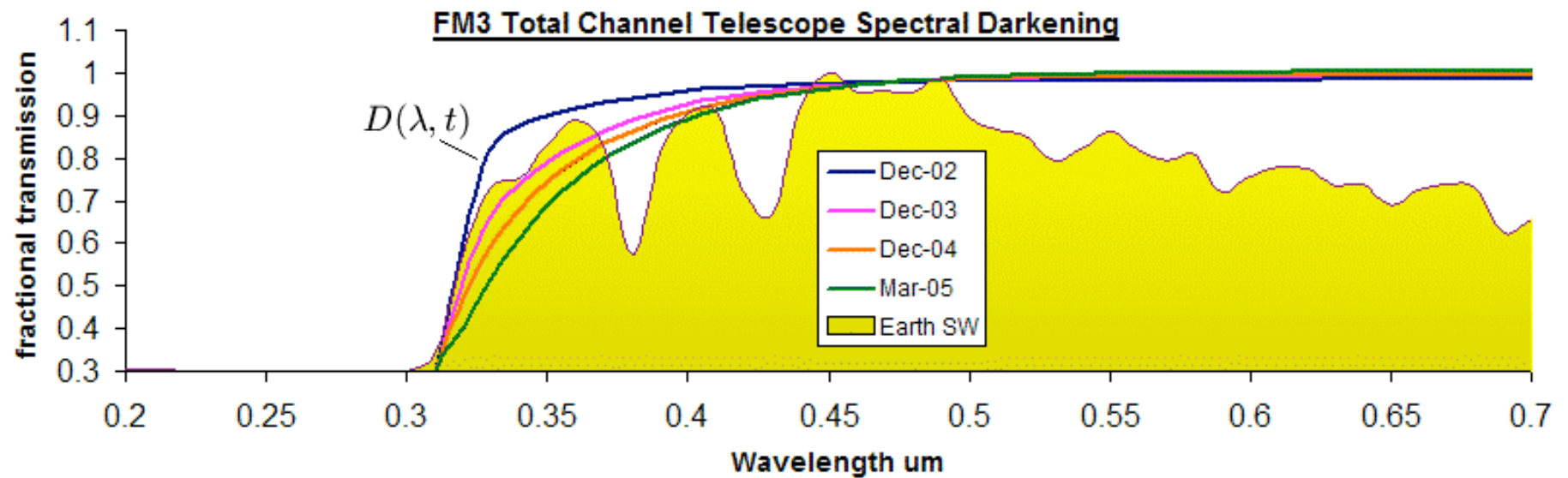






NASA Langley Research Center / Science Directorate

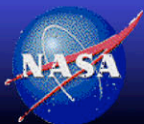




NASA Langley Research Center / Science Directorate



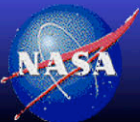
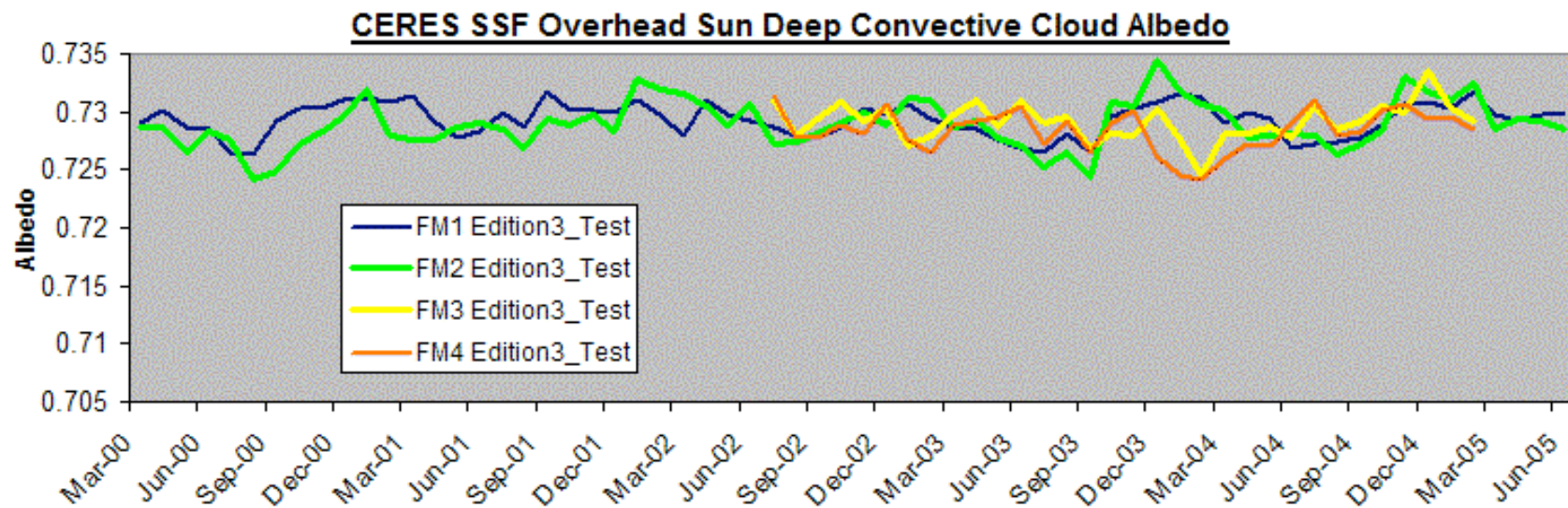
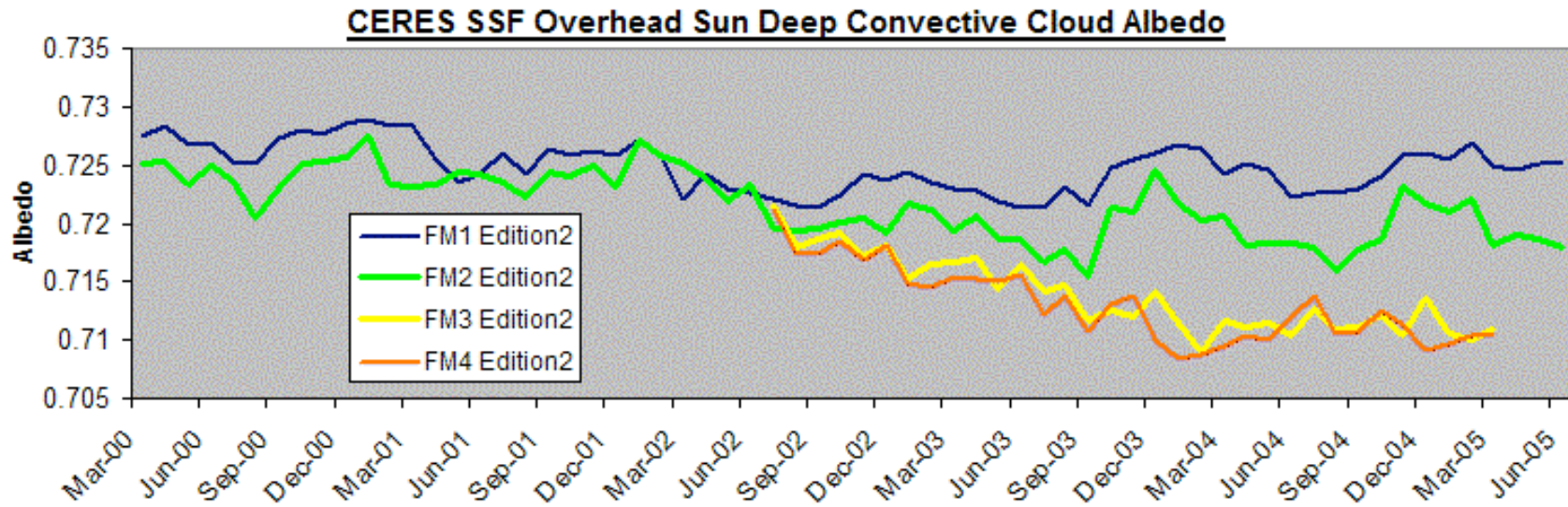
Edition 3 Test SSF data run Results **compared to Edition 2 SSF**



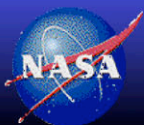
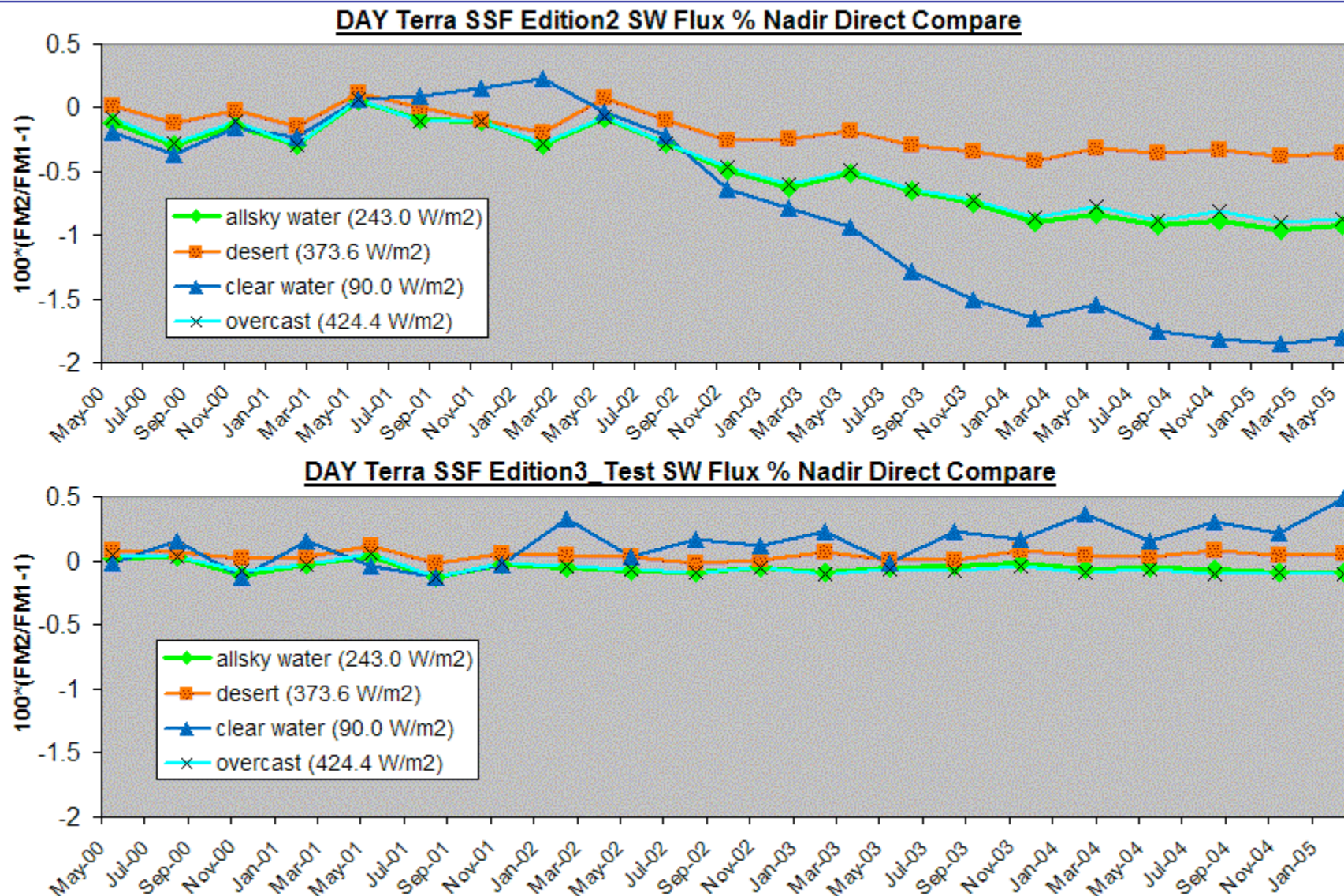
NASA Langley Research Center / Science Directorate



SSF Edition3 Test run changes from Edition2



SSF Edition3 Test run changes from Edition2

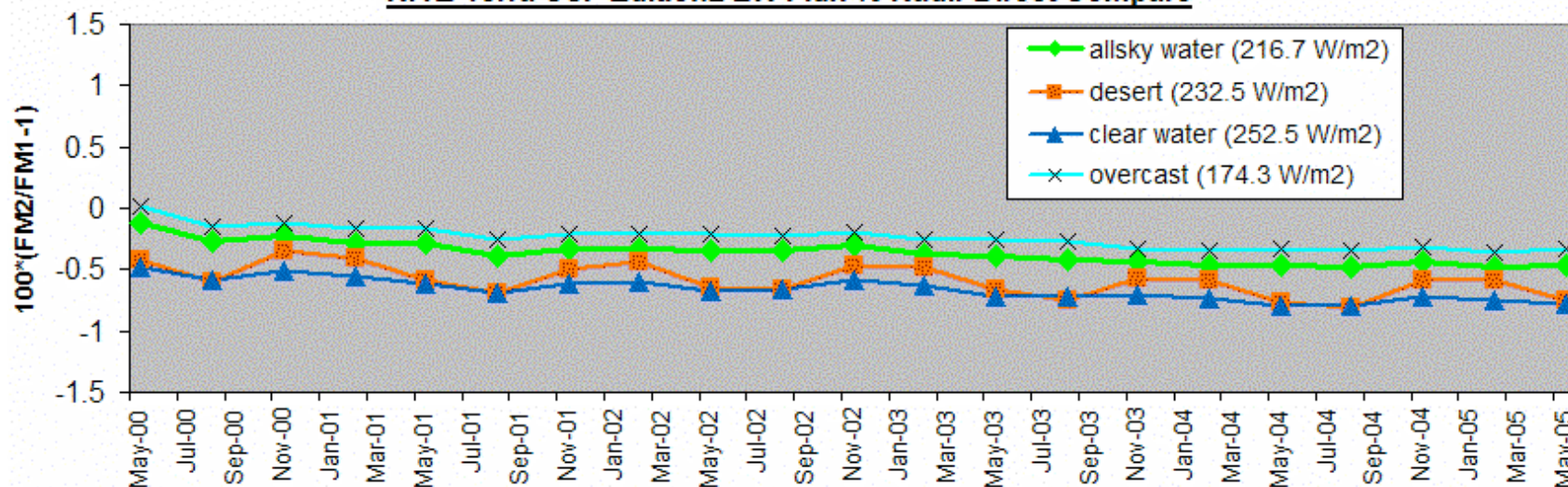


NASA Langley Research Center / Science Directorate

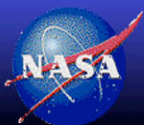
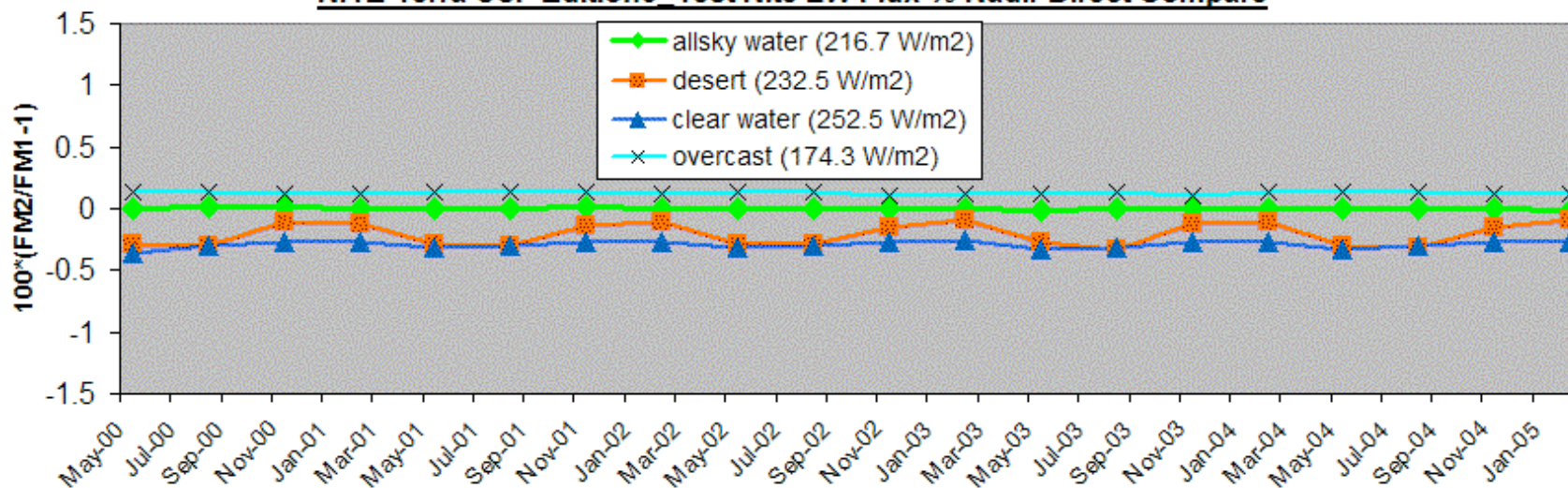


SSF Edition3 Test run changes from Edition2

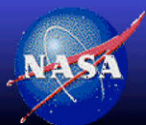
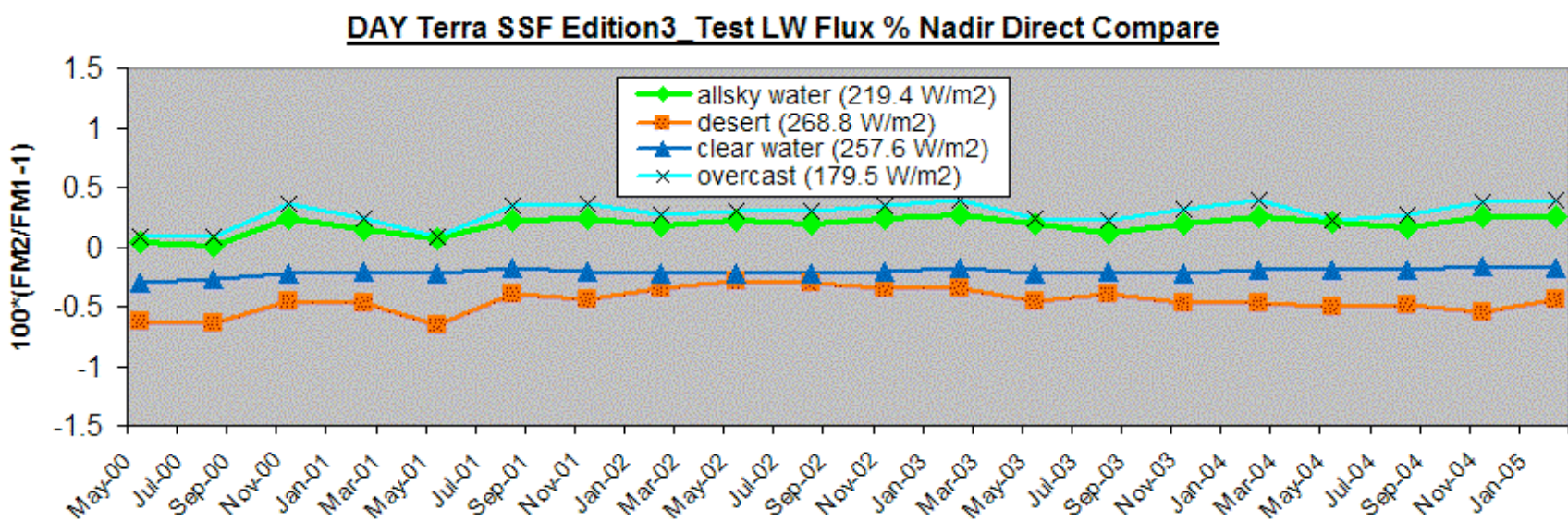
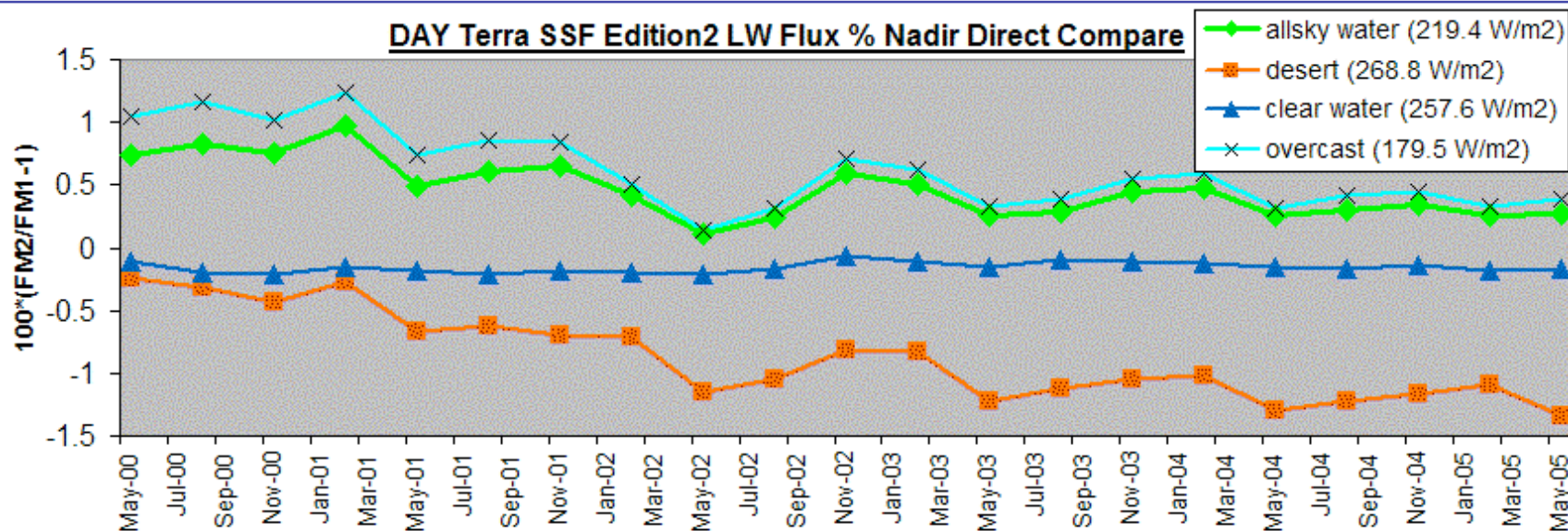
NITE Terra SSF Edition2 LW Flux % Nadir Direct Compare



NITE Terra SSF Edition3 Test Nite LW Flux % Nadir Direct Compare



SSF Edition3 Test run changes from Edition2

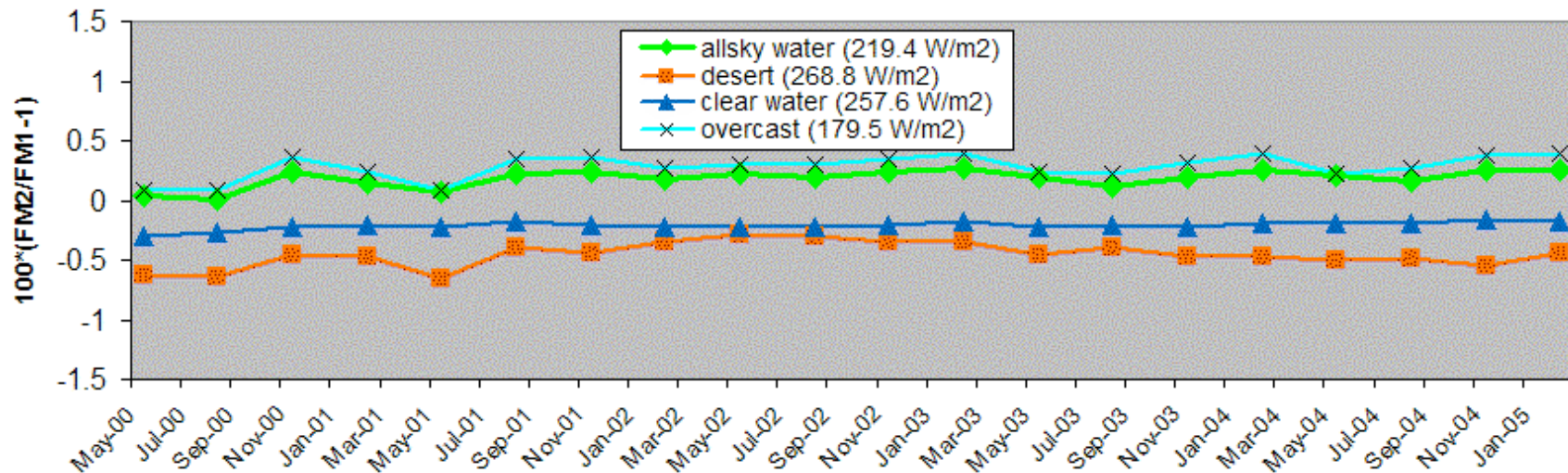


NASA Langley Research Center / Science Directorate

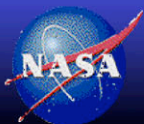
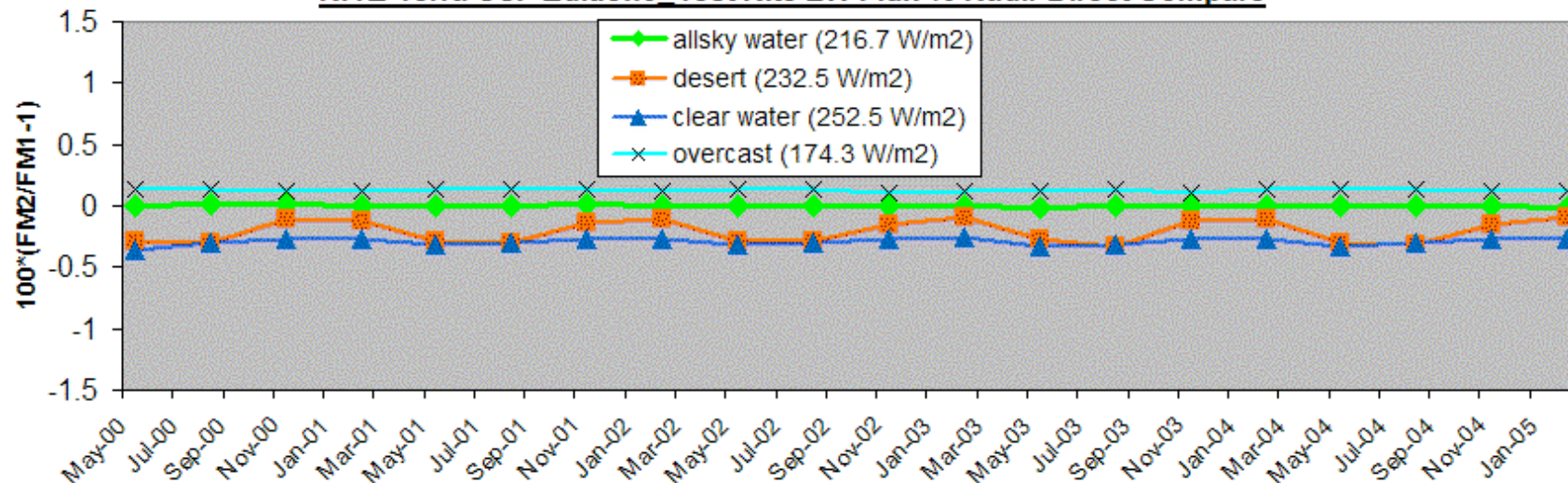


SSF Edition3 Test run changes

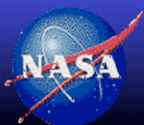
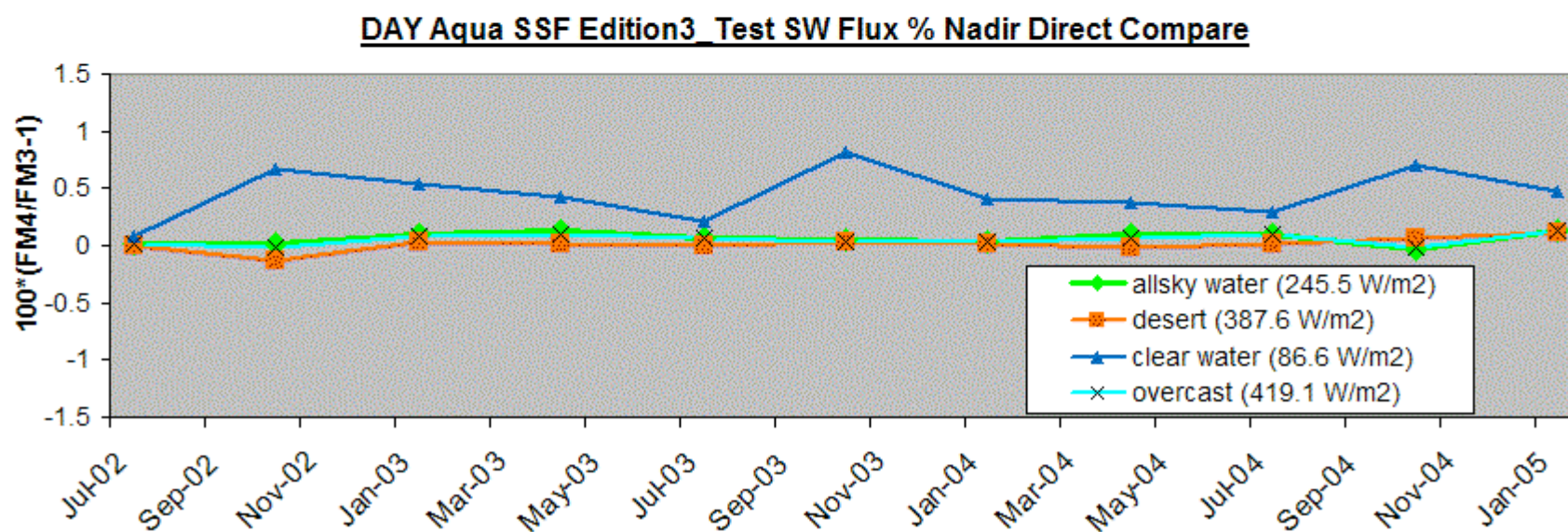
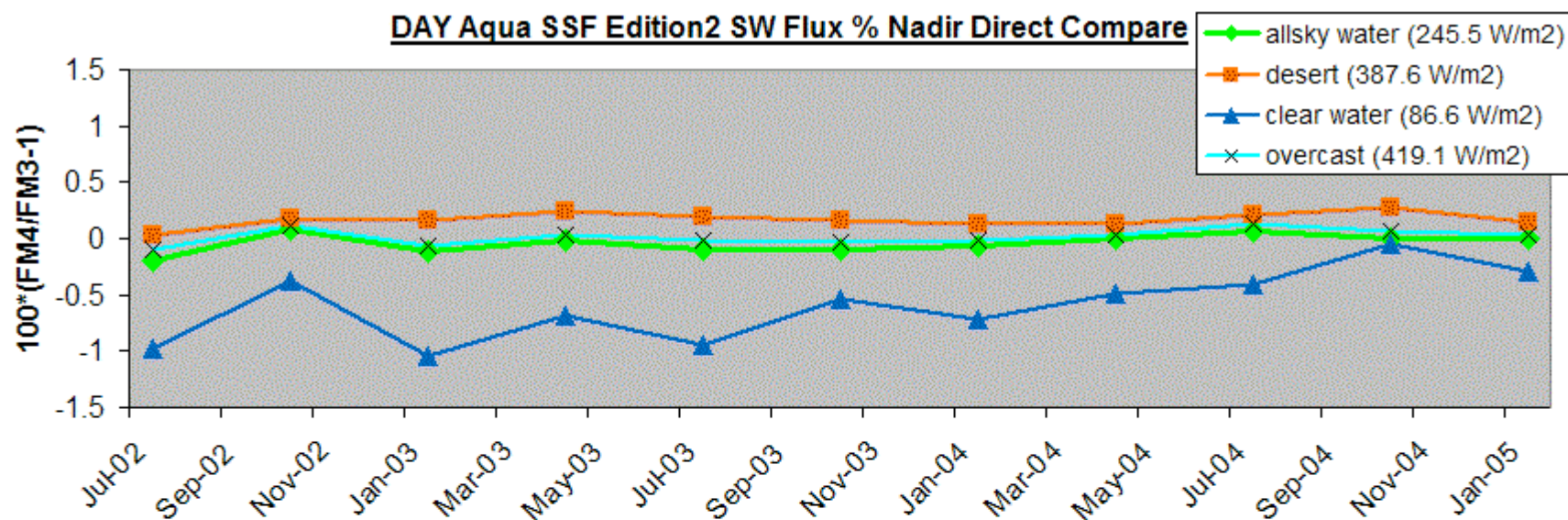
DAY Terra SSF Edition3 Test LW Flux % Nadir Direct Compare



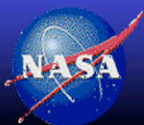
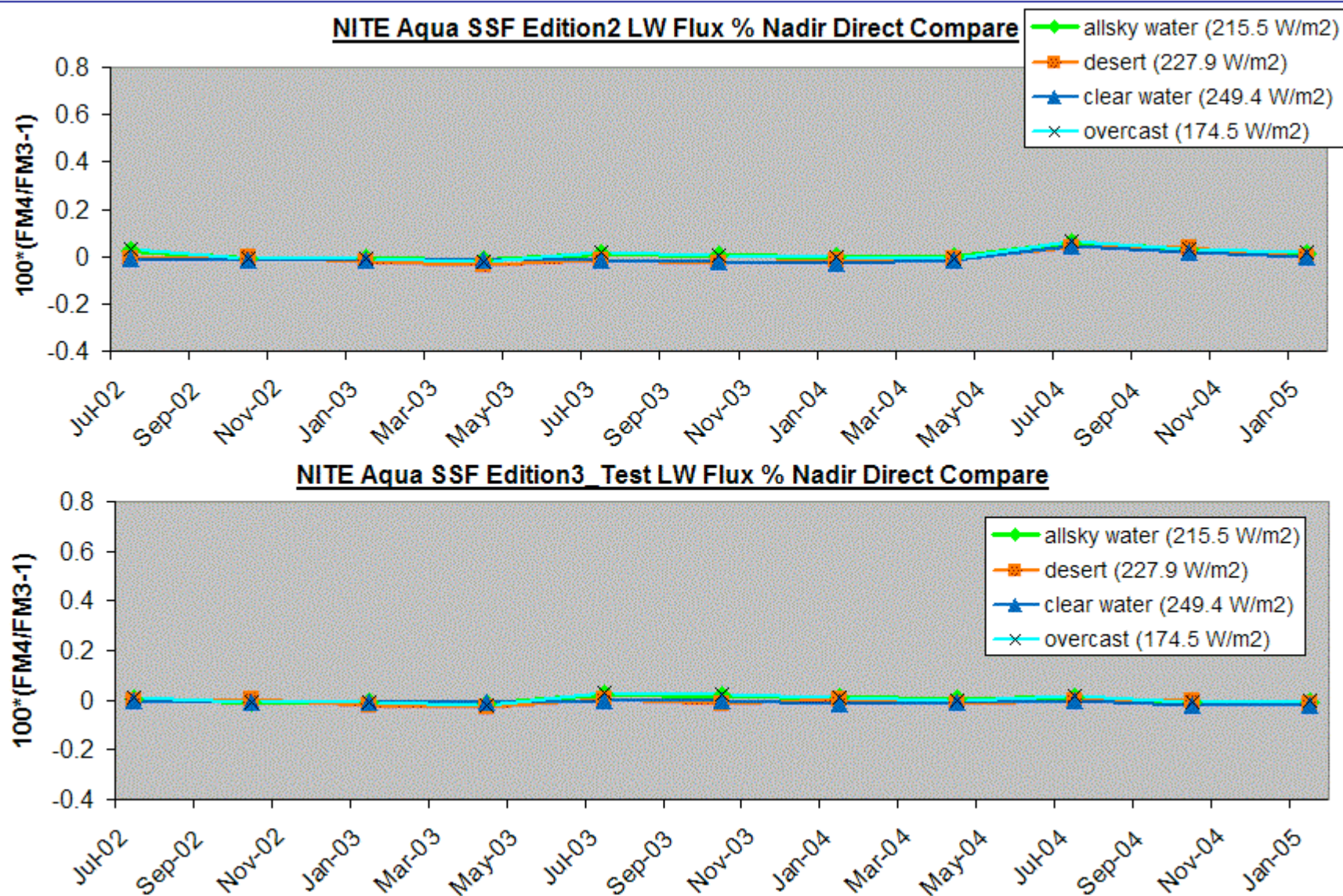
NITE Terra SSF Edition3 Test Nite LW Flux % Nadir Direct Compare



SSF Edition3 Test run changes from Edition2



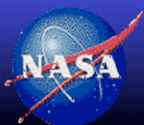
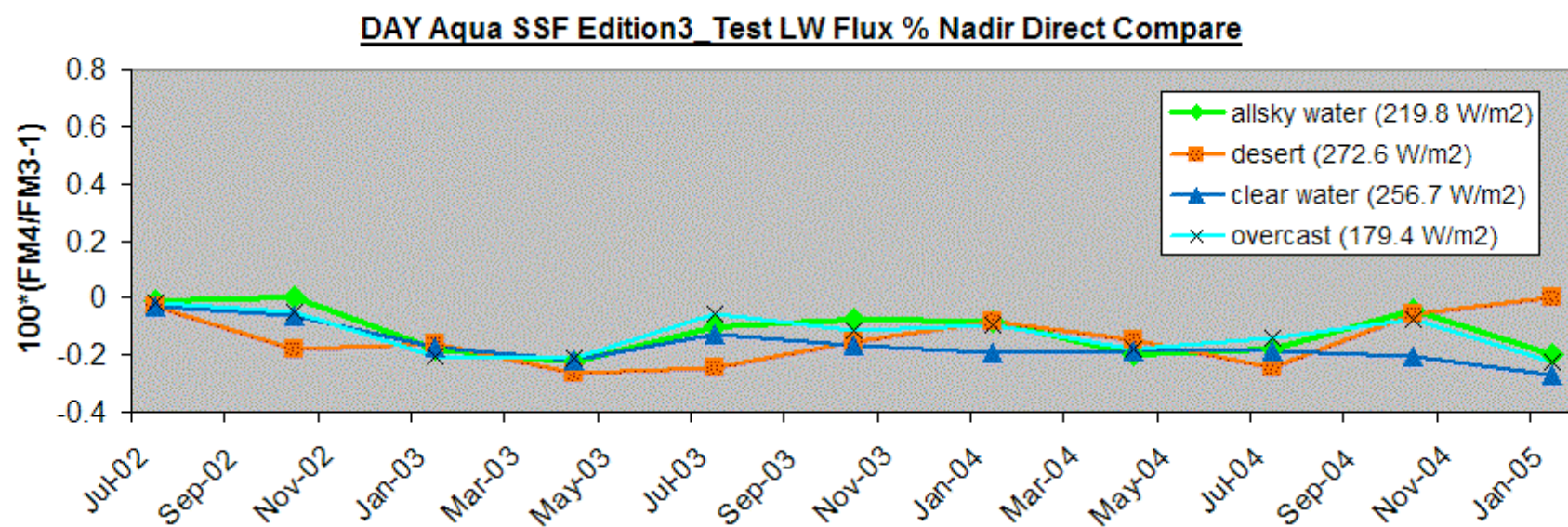
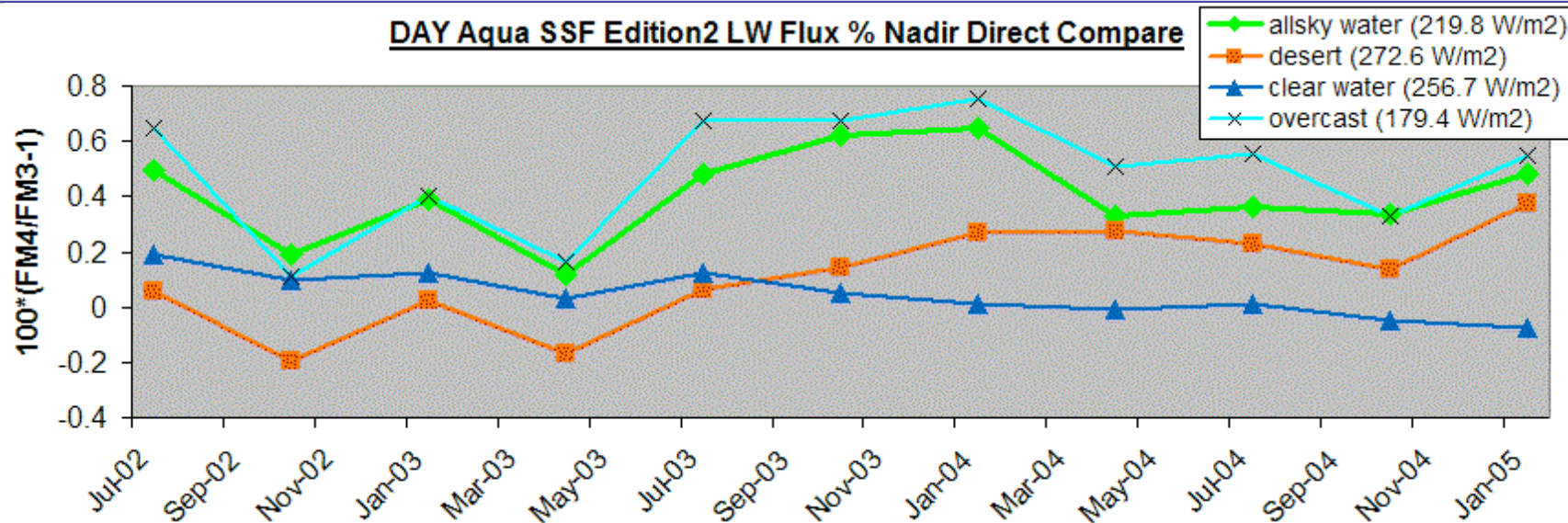
SSF Edition3 Test run changes from Edition2



NASA Langley Research Center / Science Directorate



SSF Edition3 Test run changes from Edition2

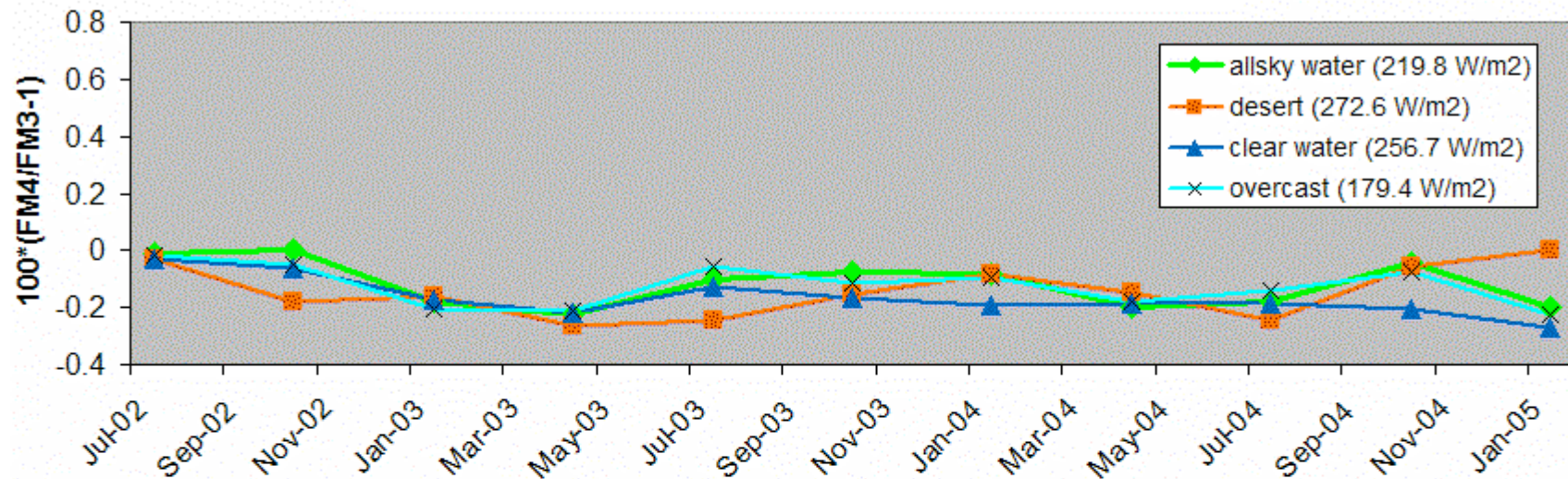


NASA Langley Research Center / Science Directorate

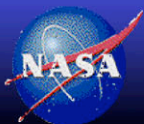
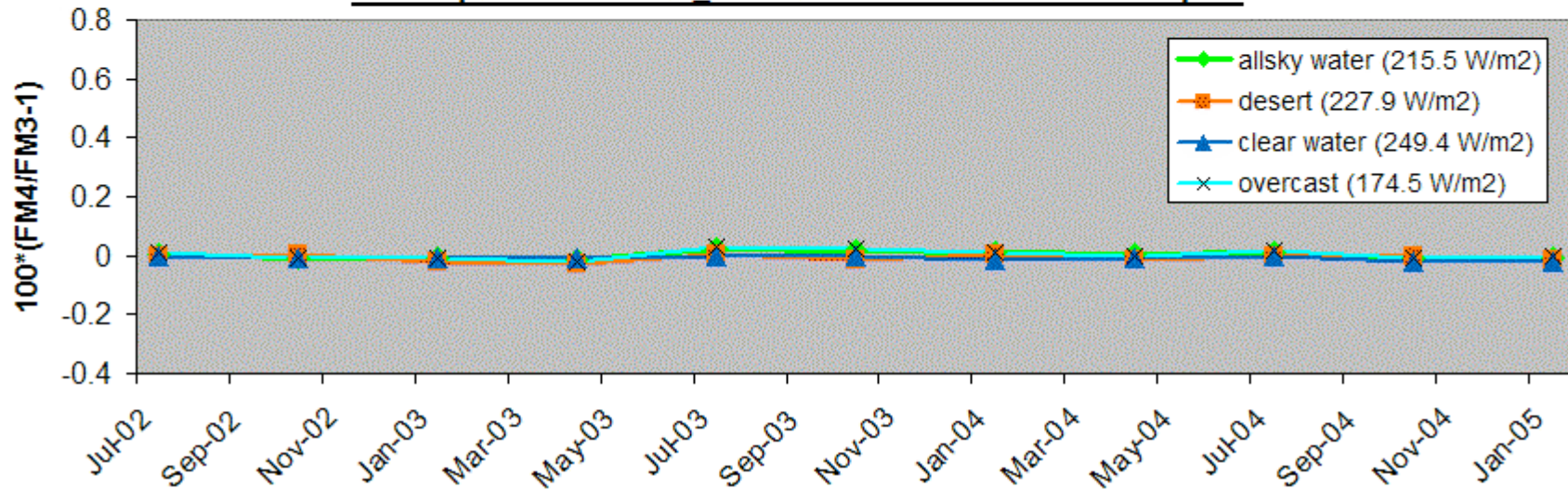


SSF Edition3 Test run changes

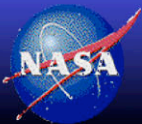
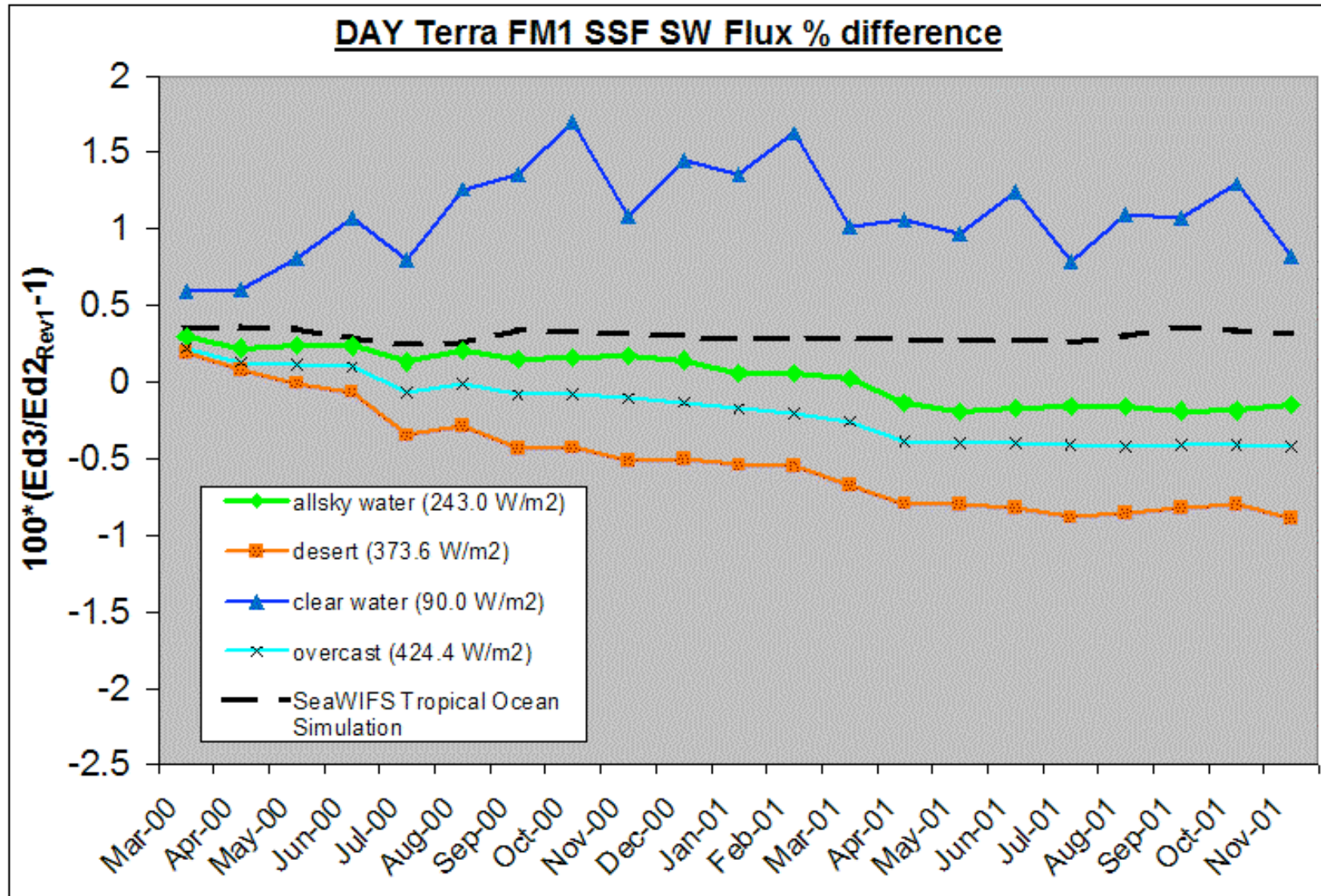
DAY Aqua SSF Edition3 Test LW Flux % Nadir Direct Compare



NITE Aqua SSF Edition3 Test LW Flux % Nadir Direct Compare



SSF Edition3 Test run changes from Edition2_{Rev1}

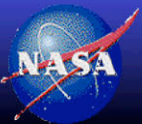
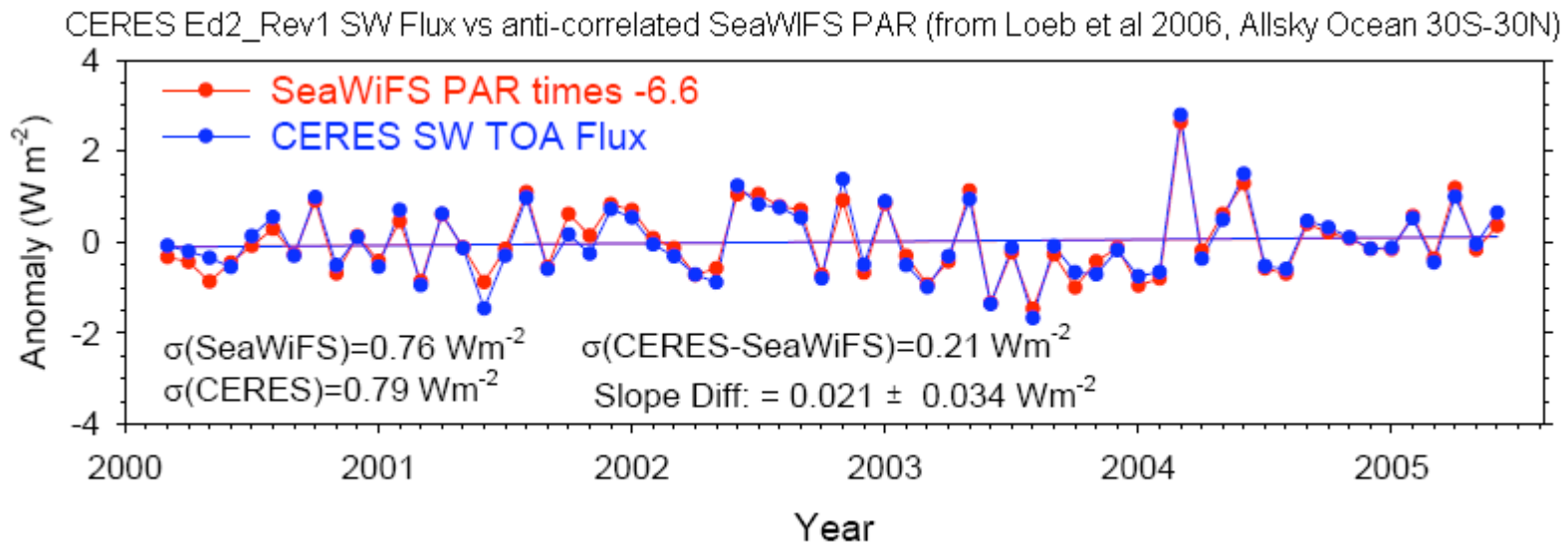


NASA Langley Research Center / Science Directorate



SSF Edition3 Test run changes from Edition2_{Rev1}

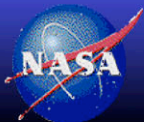
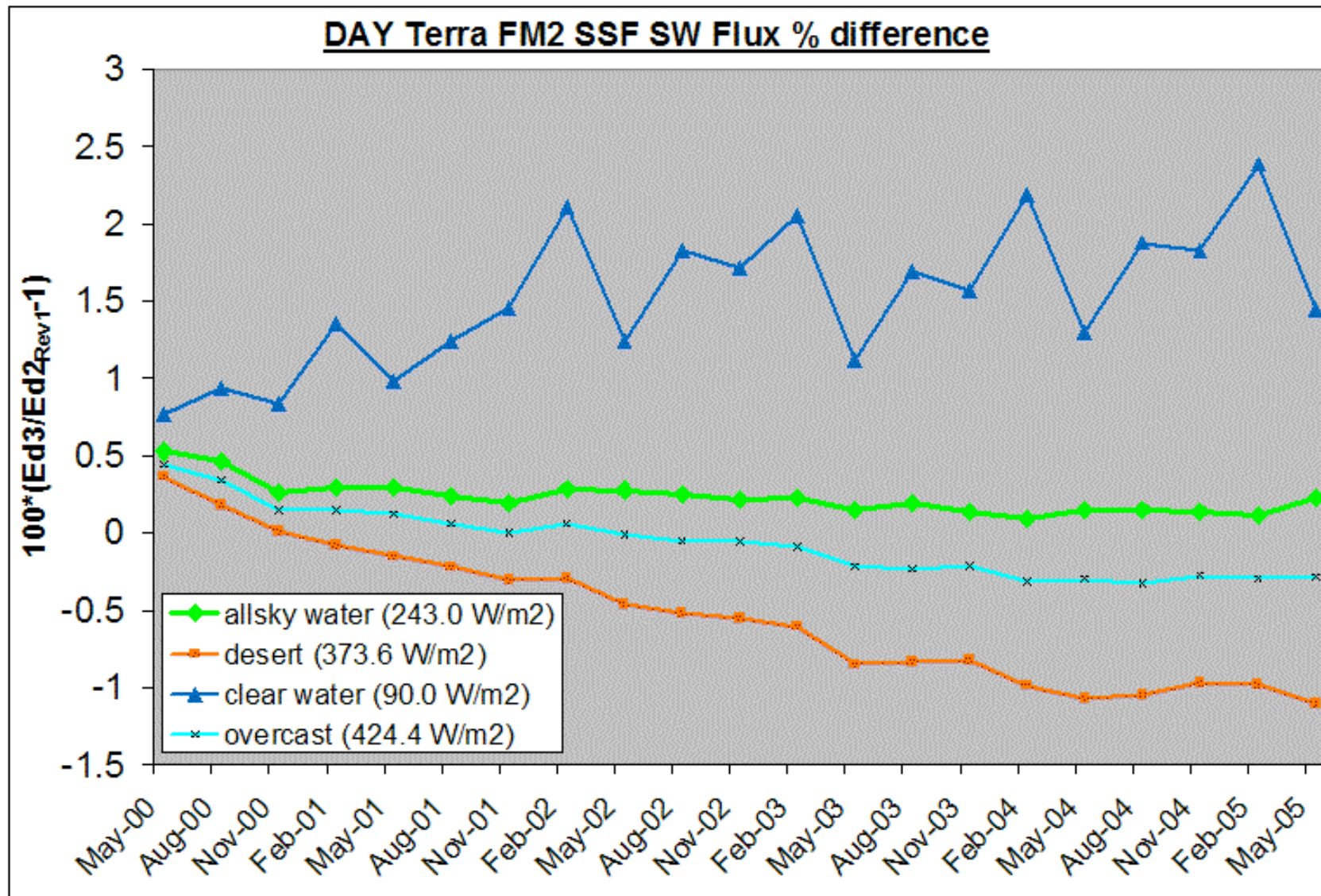
Hence SSF Edition3 Test run data should show the same excellent agreement with anti-correlated SeaWiFS PAR as Edition 2 Rev1:



NASA Langley Research Center / Science Directorate



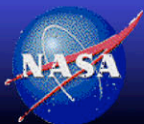
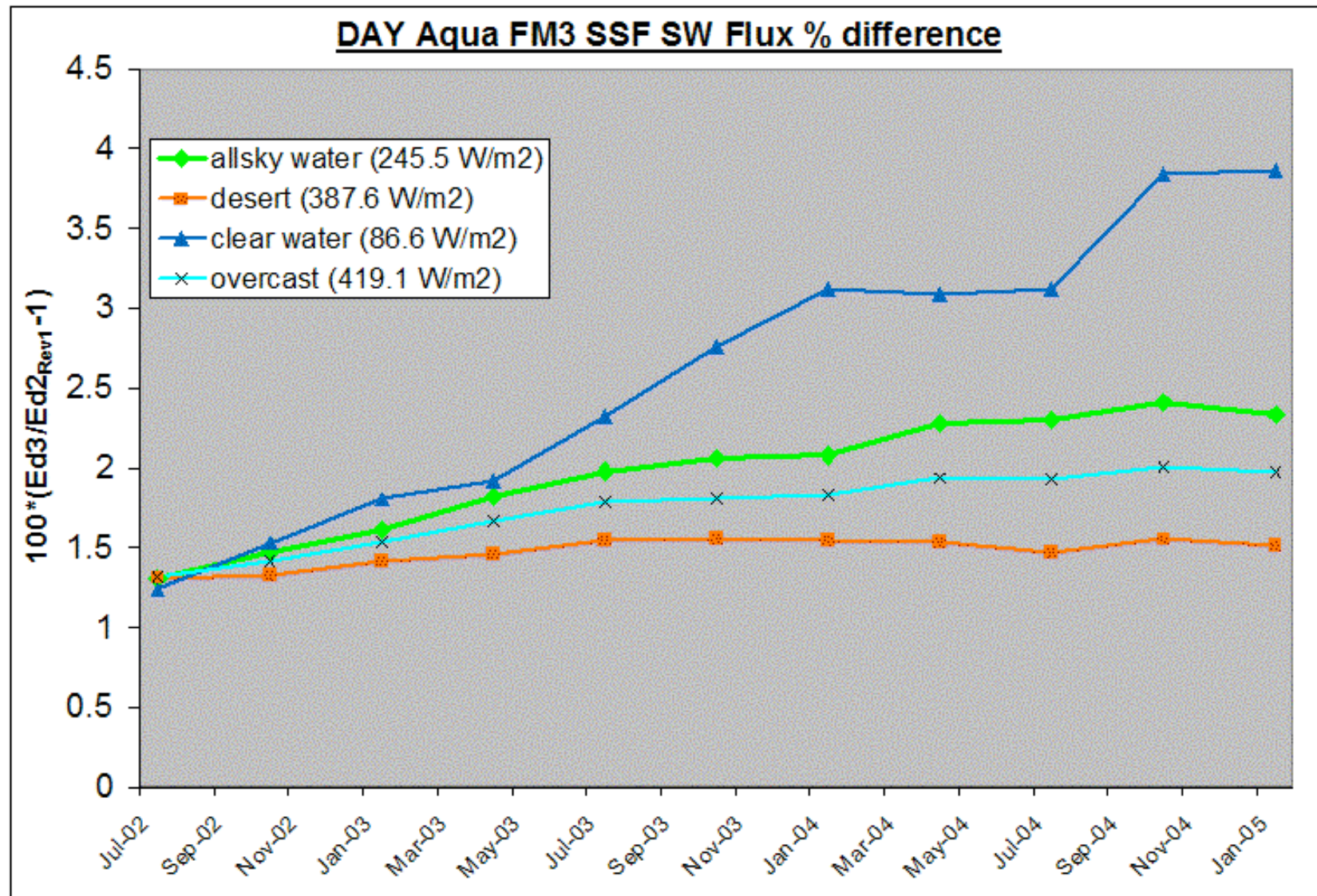
SSF Edition3 Test run changes from Edition2_{Rev1}



NASA Langley Research Center / Science Directorate



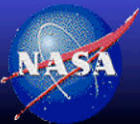
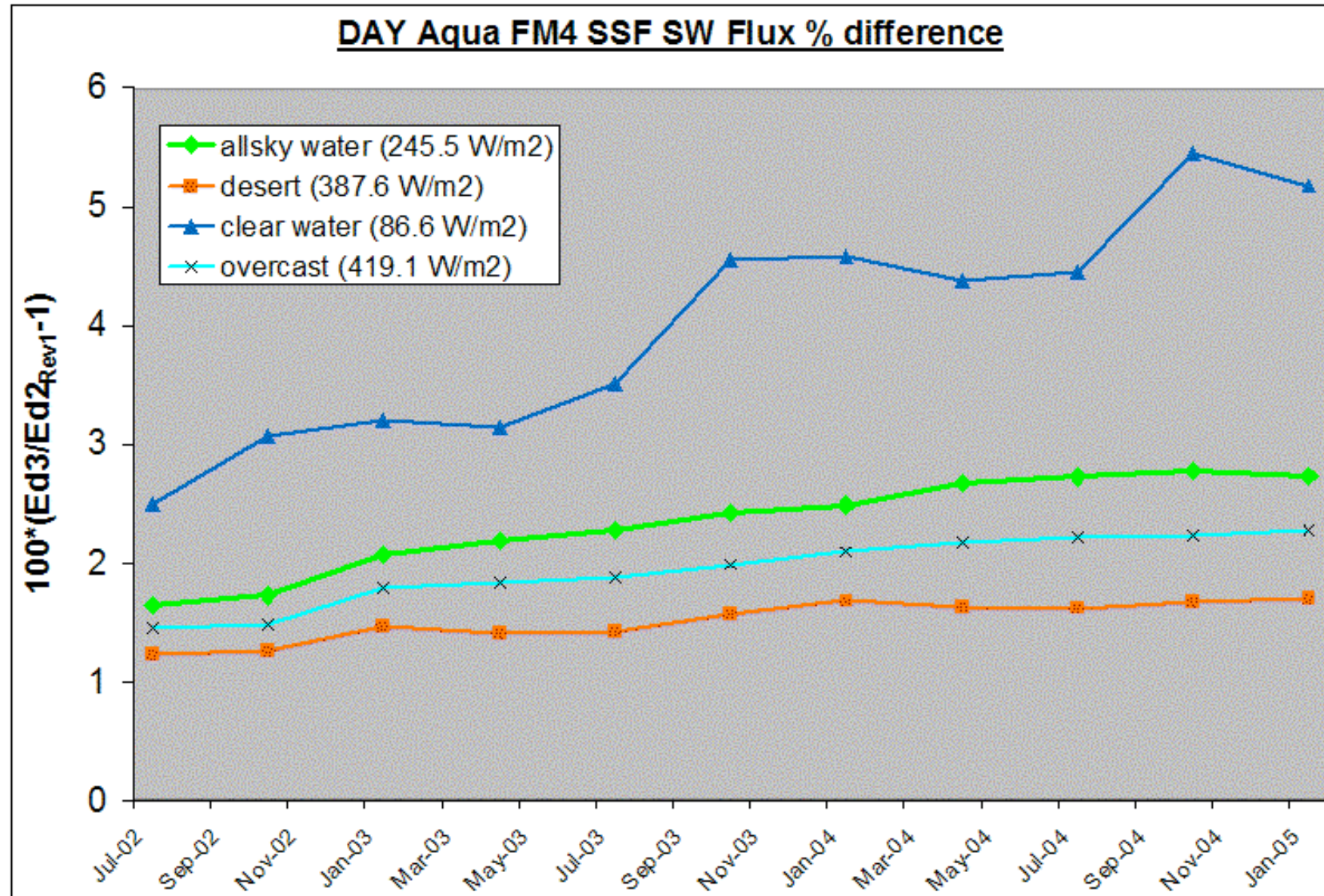
SSF Edition3 Test run changes from Edition2_{Rev1}



NASA Langley Research Center / Science Directorate



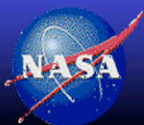
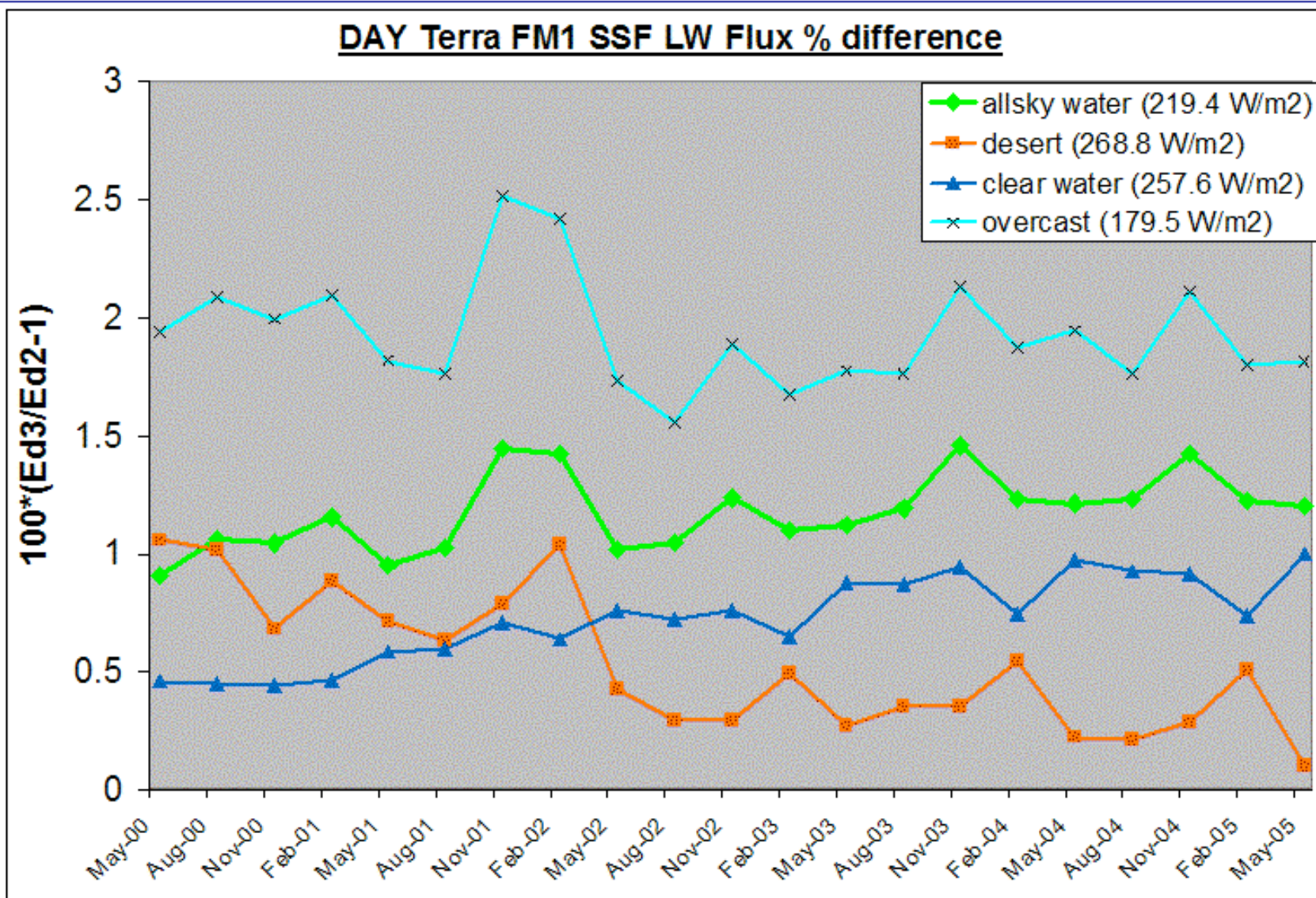
SSF Edition3 Test run changes from Edition2_{Rev1}



NASA Langley Research Center / Science Directorate



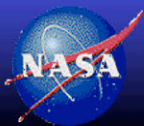
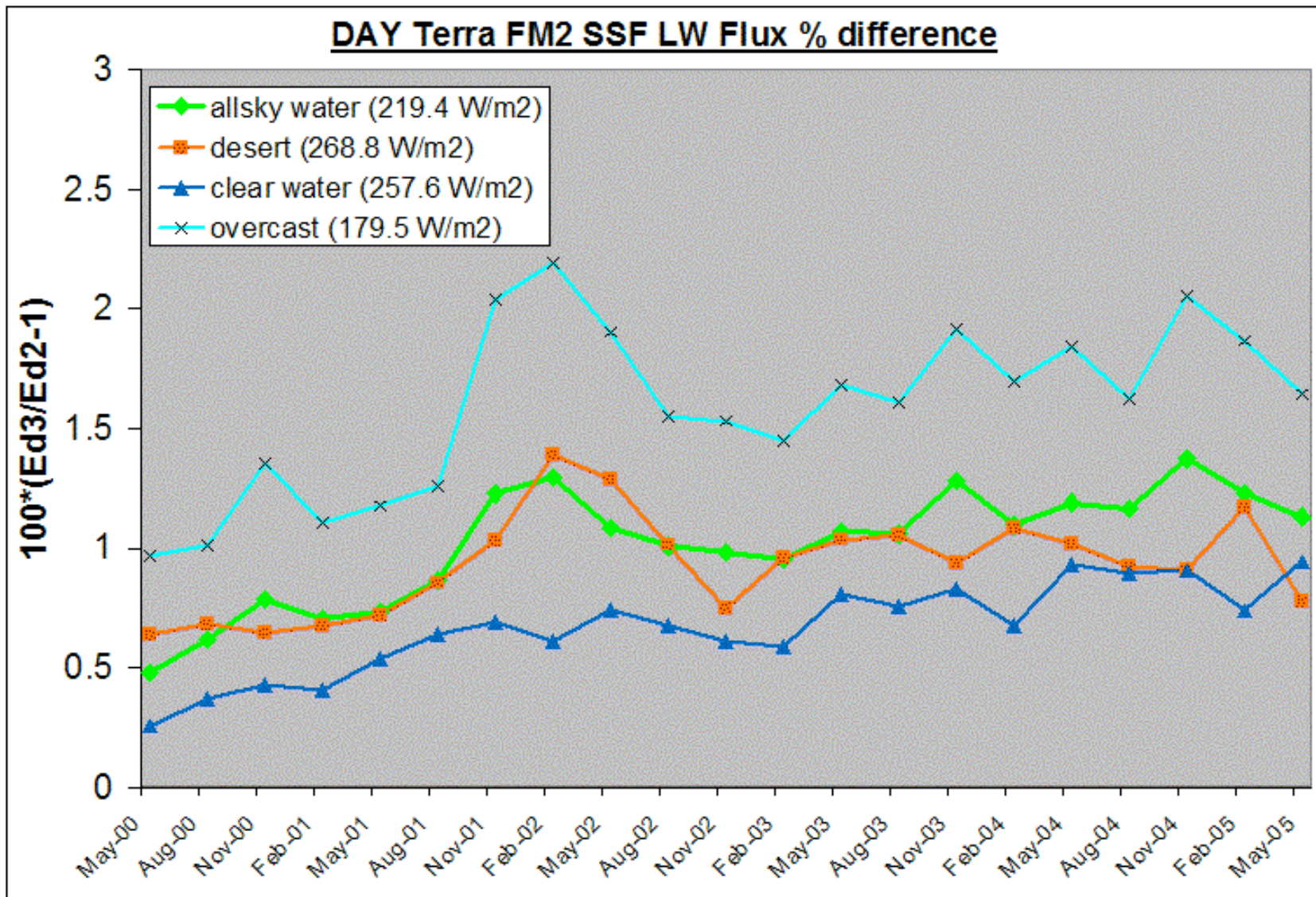
SSF Edition3 Test run changes from Edition2



NASA Langley Research Center / Science Directorate



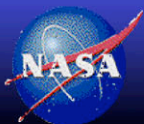
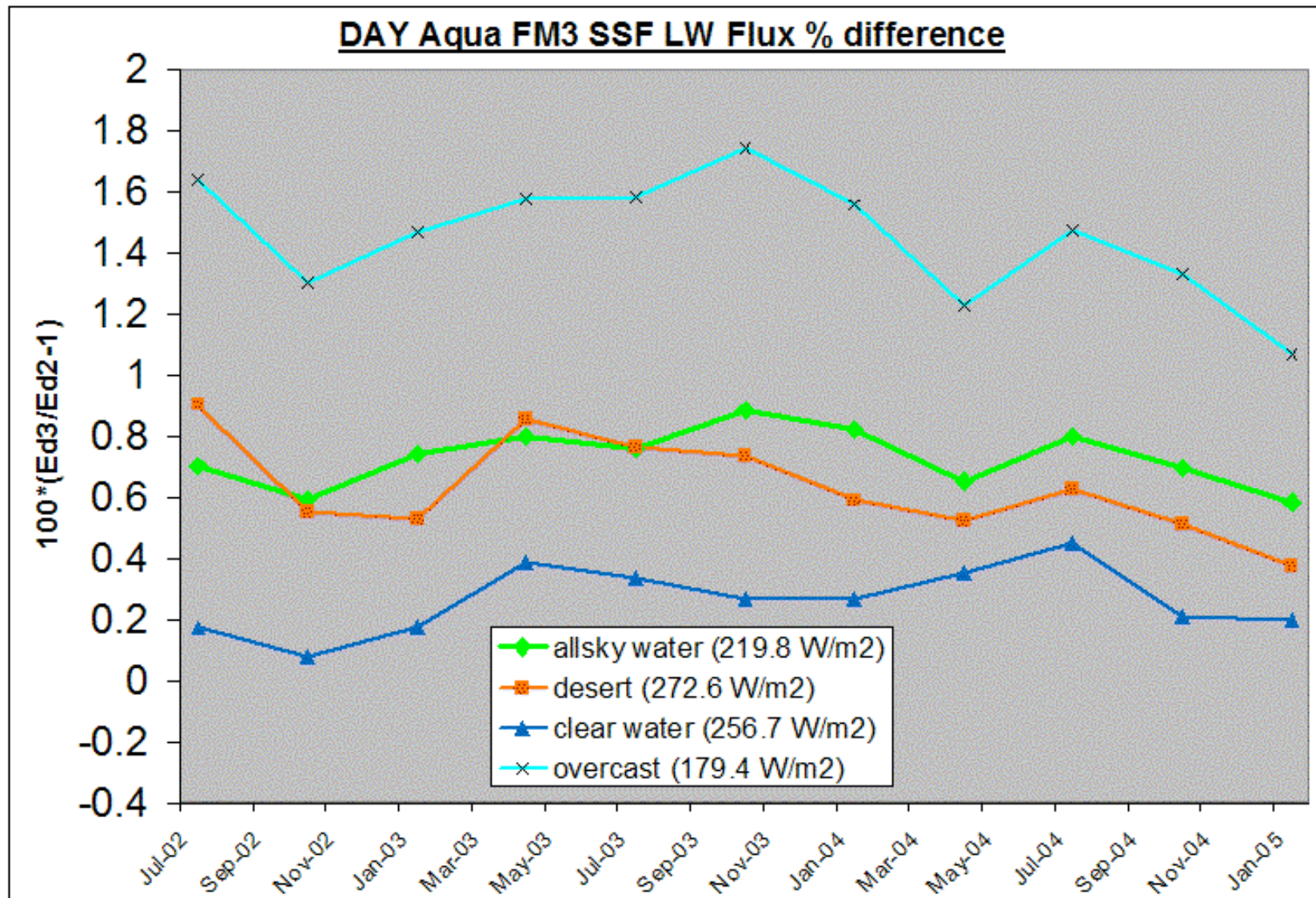
SSF Edition3 Test run changes from Edition2



NASA Langley Research Center / Science Directorate



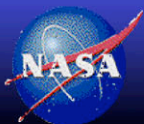
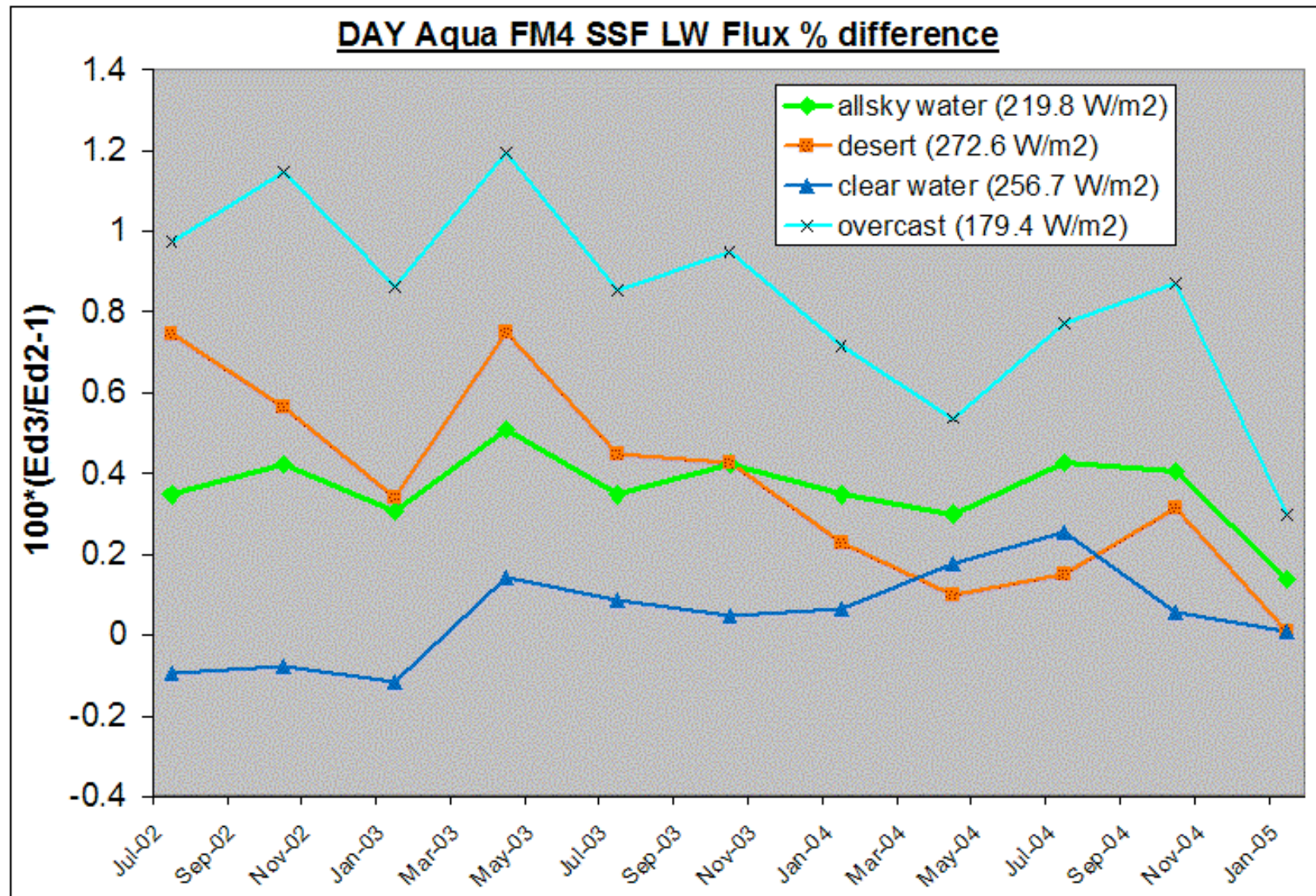
SSF Edition3 Test run changes from Edition2



NASA Langley Research Center / Science Directorate



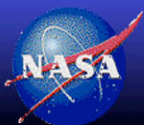
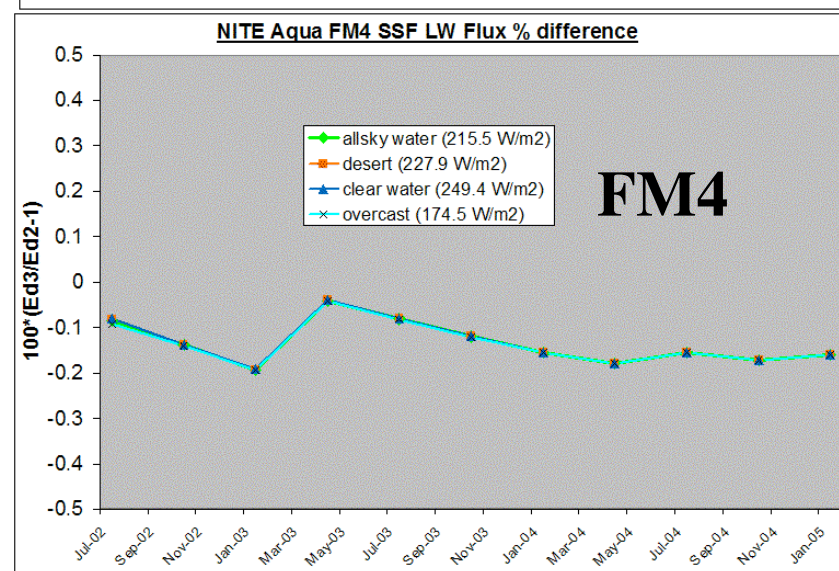
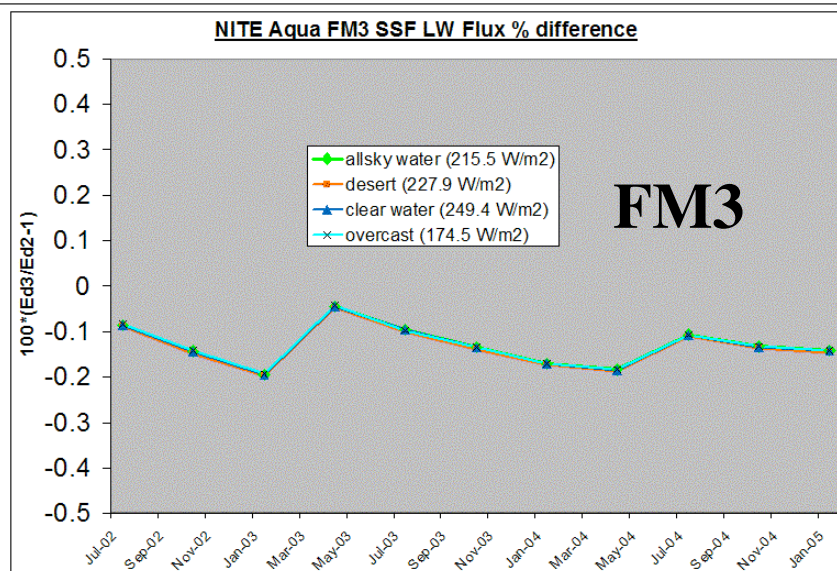
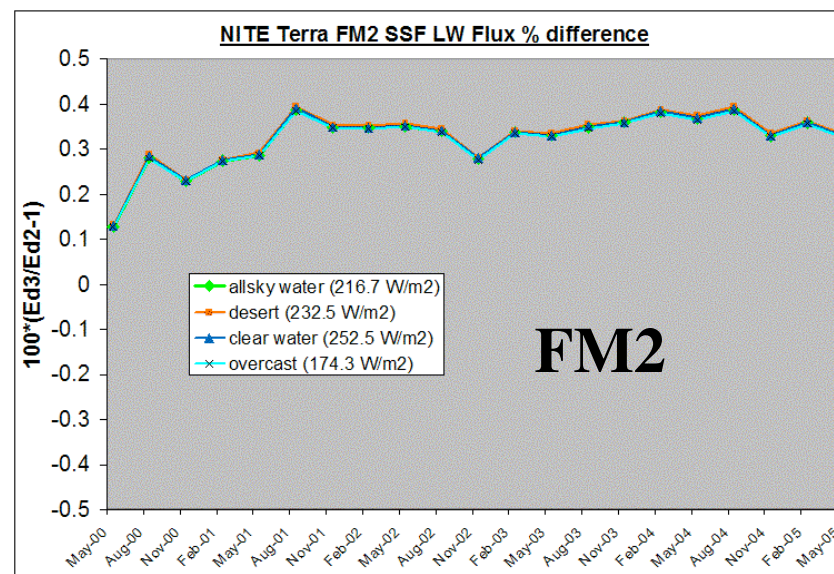
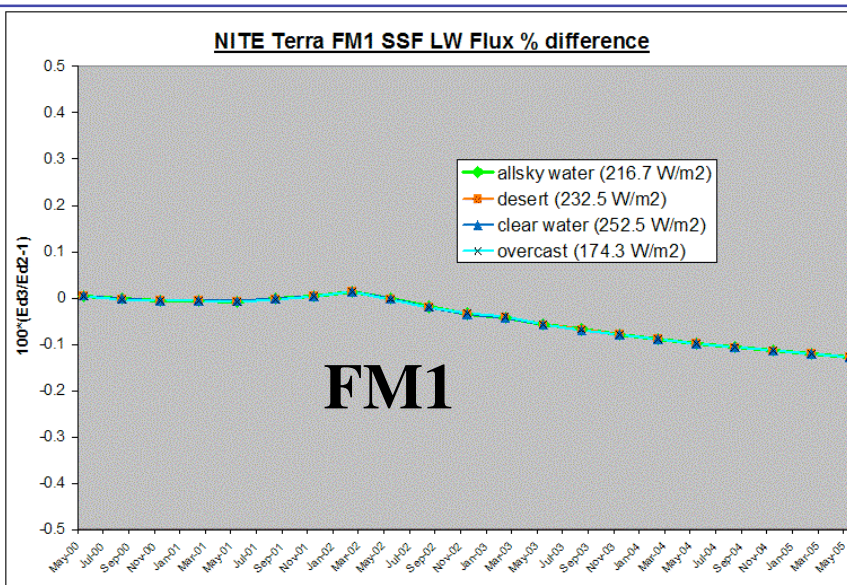
SSF Edition3 Test run changes from Edition2



NASA Langley Research Center / Science Directorate



NITE LW SSF Edition3 Test run changes from Edition2

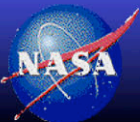


NASA Langley Research Center / Science Directorate



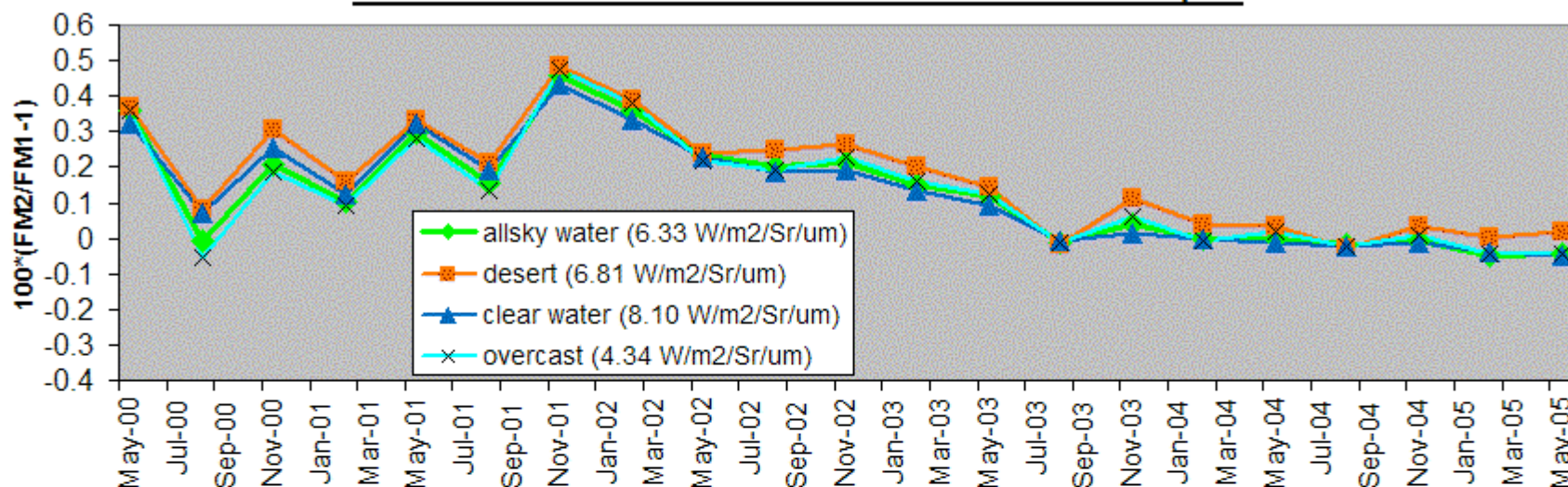
Summary & Thoughts for Comment

- Improved SW contamination spectral darkening model now produces realistic SW gain and spectral response changes for both Terra and Aqua
- These changes result in the removal of Edition 2 trends in DCC albedo from the Edition 3 test run. **Can DCC albedo change?, SeaWIFS check?**
- They also result in significant reduction in direct compare trends and scene dispersion compared Edition 2
- Terra absolute trend changes comparable to Rev1 within 0.5% while **Aqua Rev1 adjustment appears to be a too small by > 1% ?**
- Terra SW flux increase of +0.4% due to SWICS re-analysis and **using FM1 as reference standard for FM2?**
- Aqua SW flux increase of > 1% using DCC albedo, likely caused by start of mission contamination issues (**FM4 clear ocean +2.5% at start!?**)
- Daytime LW increase: Terra +1% at start and continues to increase by 0.5%, Aqua +0.6% at start with minimal allsky trend
- **Comments? Instrument Working group session for more detail.**

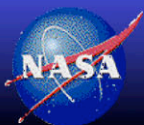
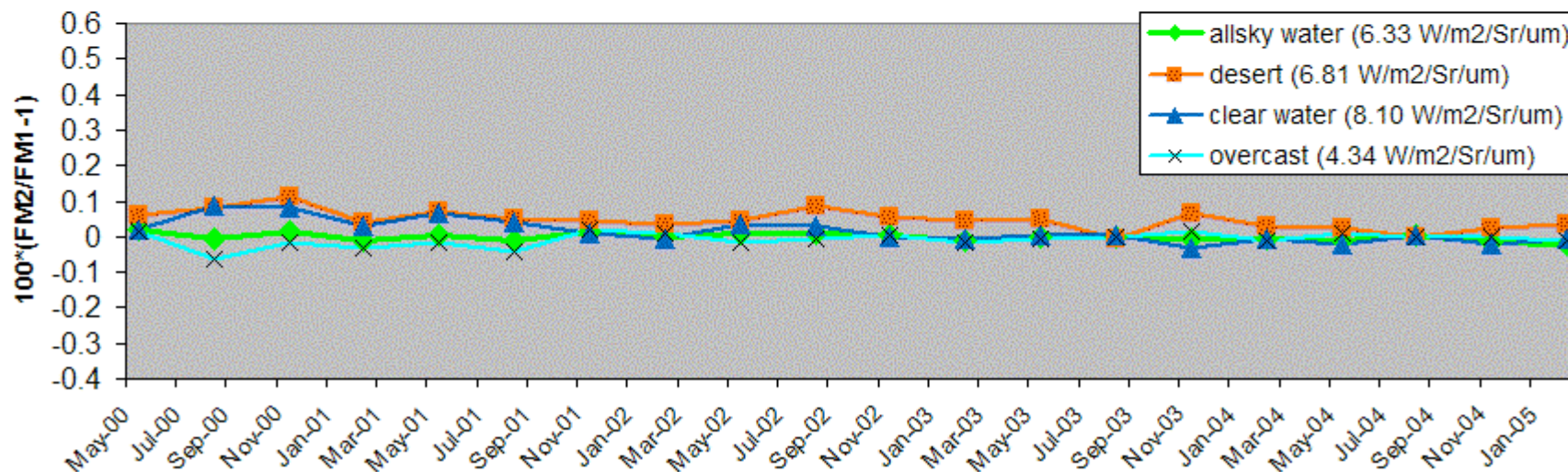


SSF Edition3 Test run changes from Edition2

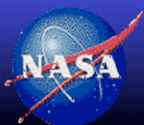
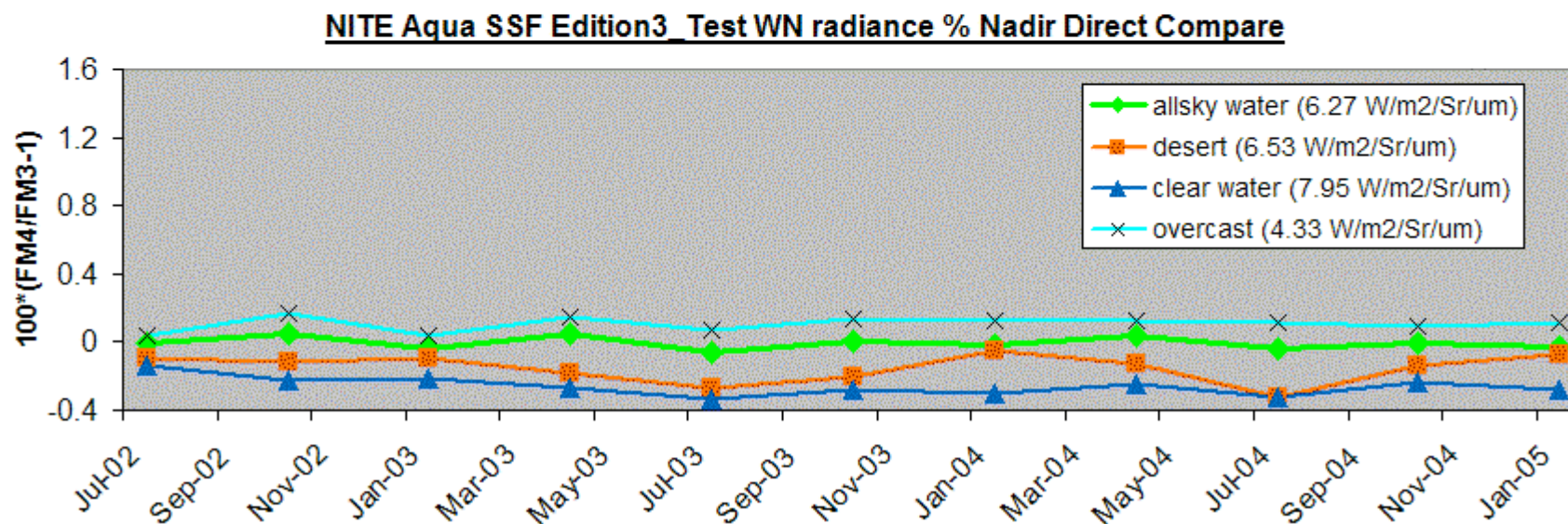
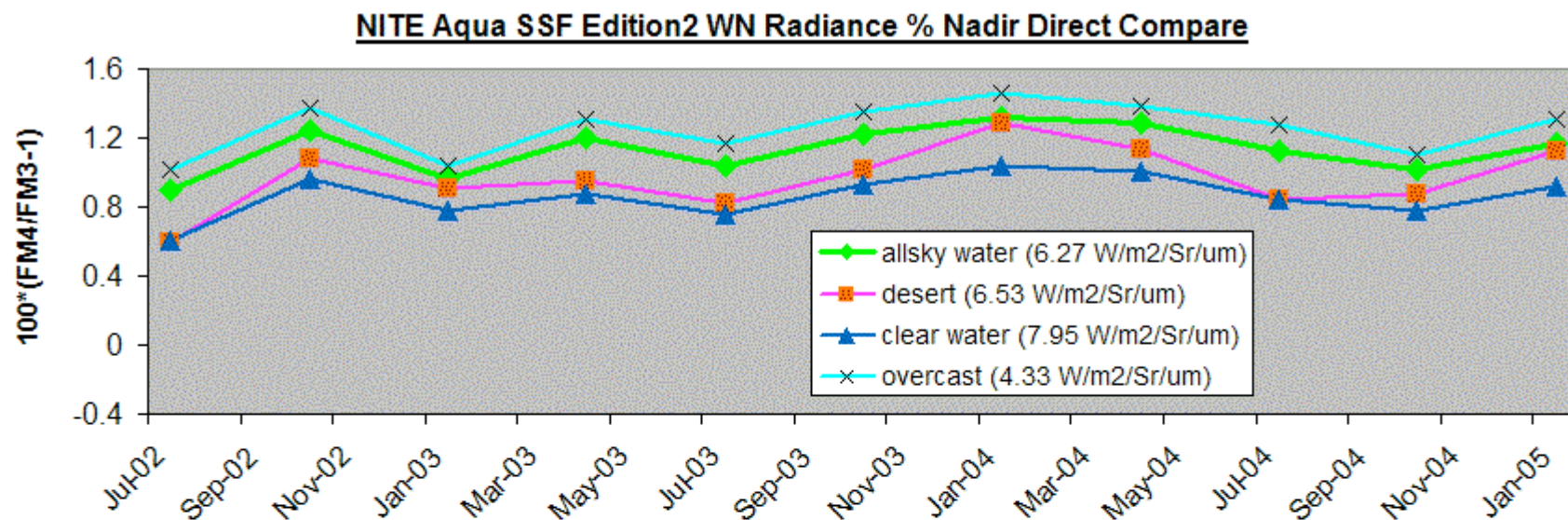
NITE Terra SSF Edition2 WN Radiance % Nadir Direct Compare



NITE Terra SSF Edition3 Test WN Radiance % Nadir Direct Compare



SSF Edition3 Test run changes from Edition2



NASA Langley Research Center / Science Directorate

